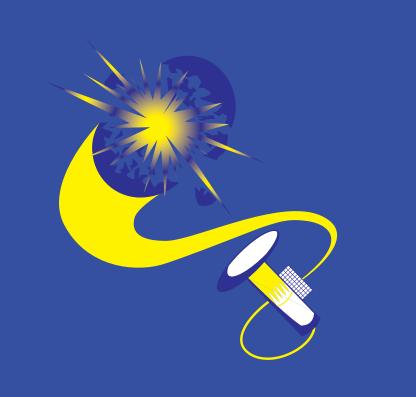


Same-Shot X-Ray Thomson Scattering and Streaked Imaging of Xenon Radiative Shock Experiments

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Creating Radiative Shocks

In a radiative shock the density and temperature structure of the shock have been altered by radiation transport from the shockheated matter.

The radiative regime is reached when radiation fluxes exceed material fluxes, i.e.:

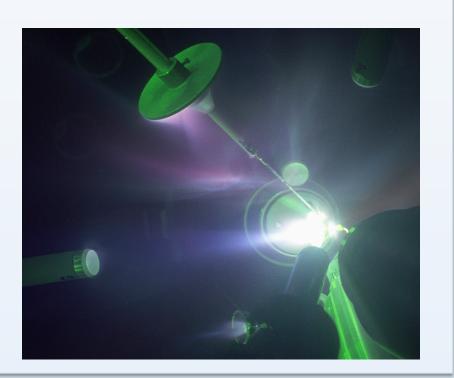
$$\sigma T_S^4 > \frac{\rho_o u_s^3}{2}$$
 or, for radiative shocks: $R_{rad} = \frac{\sigma T_S^4}{\rho_o u_s^3/2} > 1$

In the strong shock limit $T \propto u_S^2$ so $R_{rad} \propto \frac{u_S^3}{1}$

Systems with high velocities and low densities that can reach the radiative regime are ubiquitous in astrophysical systems, including supernovae, supernova remnants, and accretion disks.



The Vela Supernova remnant is an example of a radiative shock in astrophysics. Image from the NASA Astronomy Picture of the Day Collection

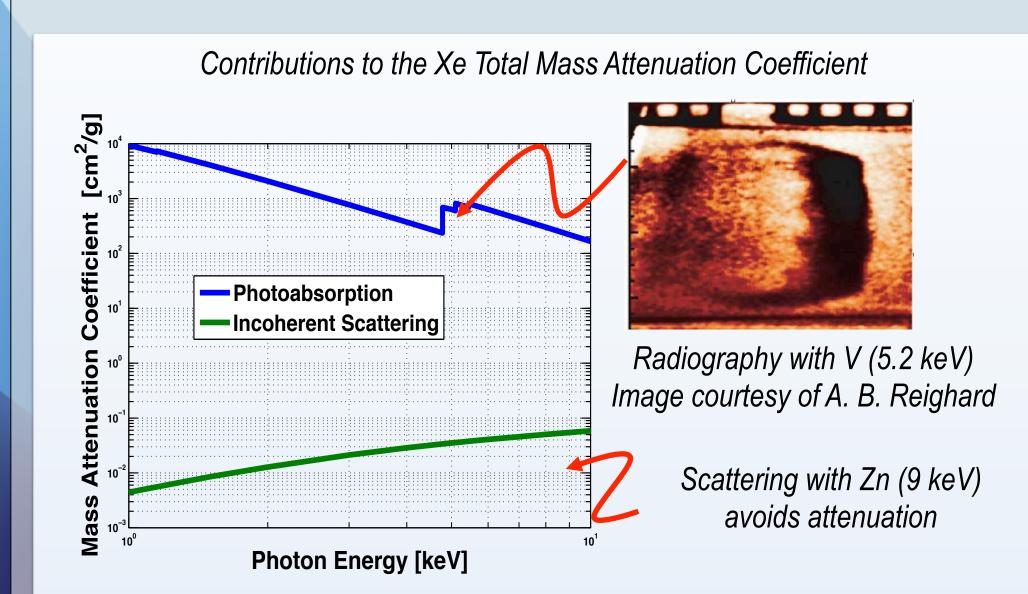


A time-integrated visible light image of the target in the chamber during laser irradiation. Plasma plumes are evident around the target and diagnostics.

Primary X-ray Diagnostics

Diagnostics are fielded to work in distinct modes:

- Passive detection: Henway spectrometer and Ross pair filters
- Transmission x-ray imaging: Streaked x-ray radiography
- Scattered x-ray spectroscopy: Zinc Thomson spectrometer
- Scattered x-ray imaging: Pinhole imaging of scattered photons
 - This is not fielded in the present experiments, but may prove to be a powerful means of measuring local electron density for dense targets and remains under develop.







Target Design for 2011 X-Ray Thomson Scattering Experiment Acrylic "scaffold" Beryllium drive disk Acrylic facilitates Gold shield 10 drive beams with SSD accurate positioning of These shield the zinc xlaunch the shock. components and serves ray sources from the to provide alignment spectrometer and are Scattering slit features Thomson Scattering Angle laser cut to precisely 150 or 400 micron slit in the gold Angle is defined by the input x-ray locate the scattering slit.. shield defines the scattering volume source and vector to the diagnostic. Polyimide shock tube Thin PI contains the gas but has minimal effect on x-rays Zinc x-ray source above a few keV ZSPEC in "Crystal Up" Geometry Collimating slit Serves to define the k-vector for 13 x 13 mm aperture scattering. 4.7 mm 1 mm Streak camera fiducial wire X-r source is ~2.8 mm from TCC: 1 mm vertical, 2.6 mm left

Diagnostic Principles

scattered photon

Incoherent Thomson scattering

incident photon

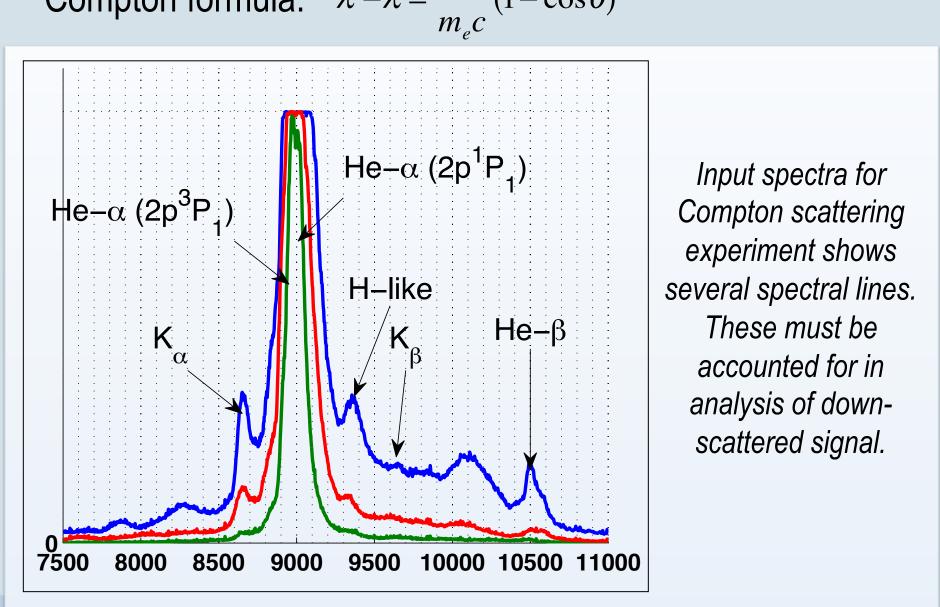
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Incoherent Thomson scattering

Involves photon scattering from individual electrons.

Recoil of free electrons leads to an energy downshift in the scattered spectrum.

 The shift is given by the Compton formula: $\lambda' - \lambda = \frac{n}{m_c c} (1 - \cos \theta)$



Streaked x-ray Radiography

Imaging a streak camera film with long-duration backlighting provides information about shock velocity and absolute position.

XRTS results

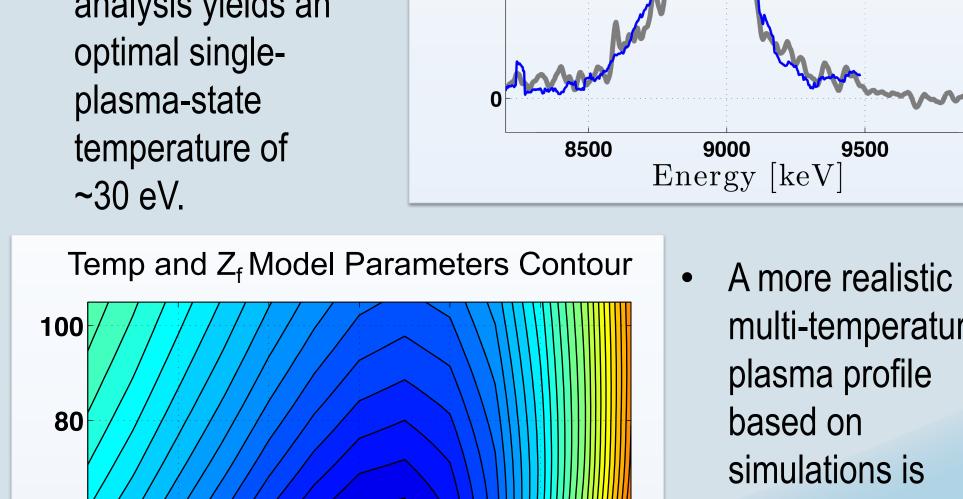
 Scattered spectra shows a clear inelastic peak, which is fit for the 400 micron-slit 400 µm XRTS Data and Fits target. Low signal-to-

Te = 30 eV

Zf = 45

noise limits the use of data from the 150 micron case.

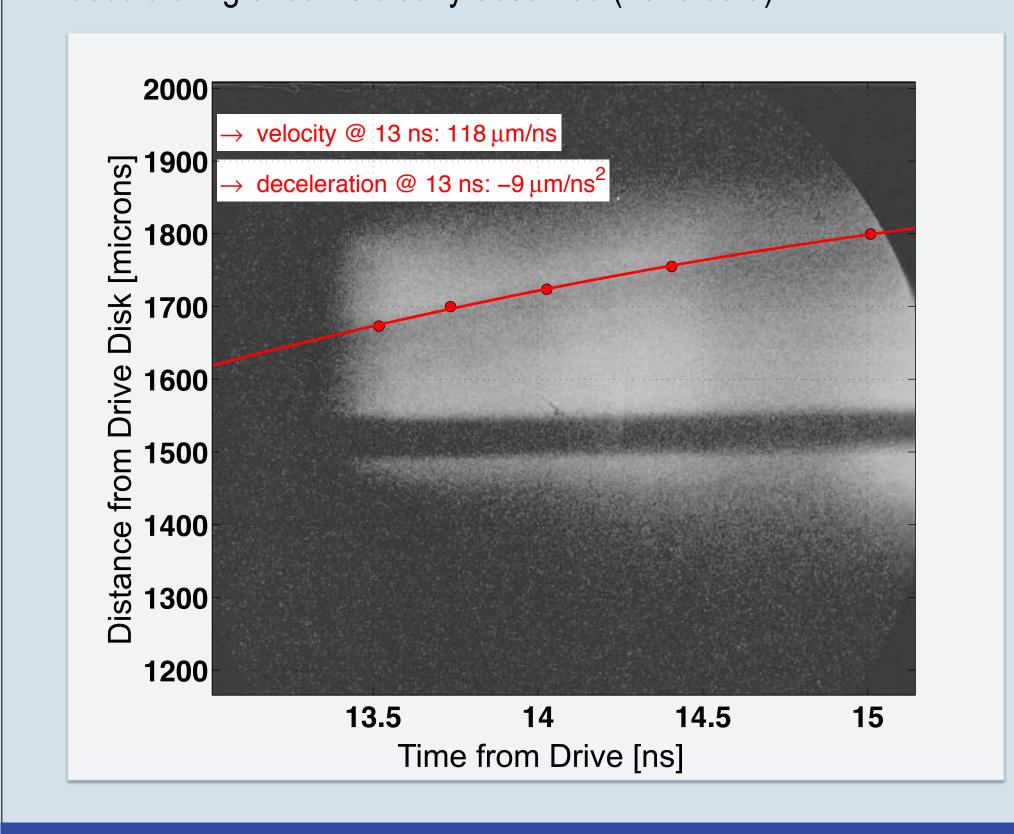
 Least-squares analysis yields an optimal singleplasma-state temperature of ~30 eV.



multi-temperature plasma profile based on simulations is ongoing, and will be verified with future high resolution experiments.

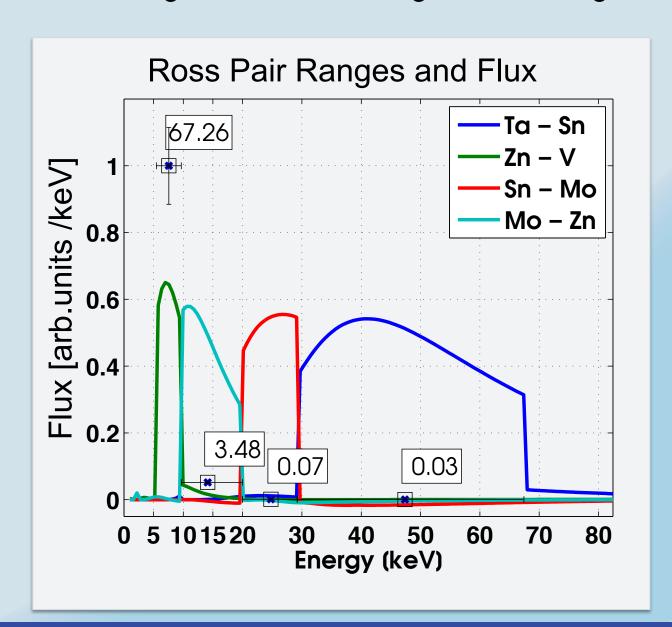
X-Ray Streak Data

- Streak camera records shock front progression at sweep speeds of 2.5 ns and 5 ns, providing time resolution ~100 ps.
- A decelerating shock is clearly observed (2010 data).



Passive X-Ray Detectors

- Passive detectors confirm x-ray lines from expected diagnostic sources and inform our understanding of the x-ray background.
- Ross pairs serve as a simple, robust diagnostic to obtain absolute photon yields.
- The Henway x-ray spectrometer consists of four independently-filtered crystals covering 2-13 keV, monitoring x-ray sources and background in the range of the diagnostics.



(Above) Elemental filters are matched in thickness and transmission to form the Ross pair matrix.

(Left) Ross pairs provide a simple, coarse spectrographic information about the xrays present during the experiment.

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