



Results of a supersonic, single-mode, shockwave-driven Kelvin-Helmholtz instability experiment

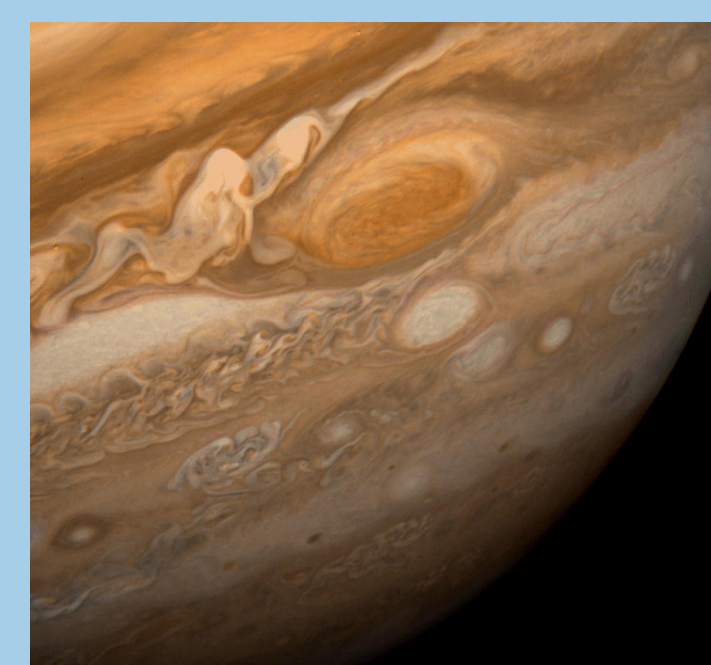
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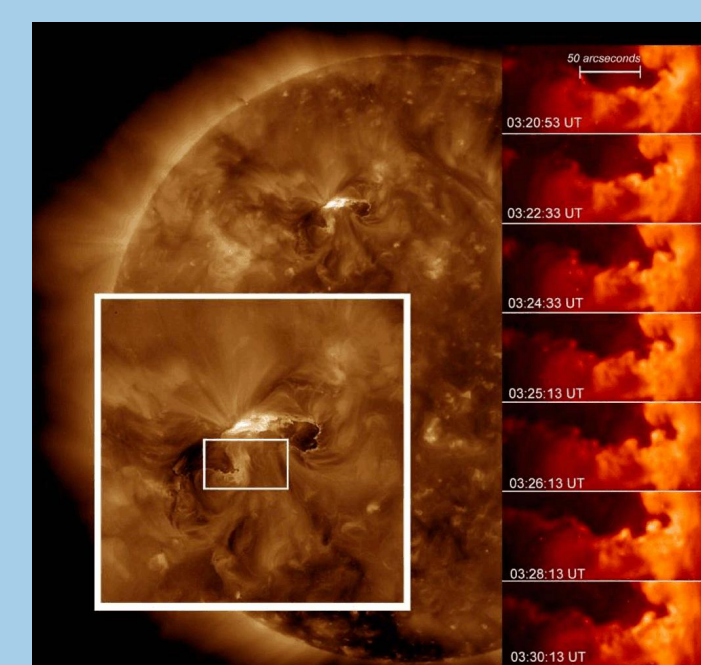
Motivation

- The Kelvin-Helmholtz instability is one of the most common hydrodynamic instabilities, and is prominent in astrophysics and fusion.

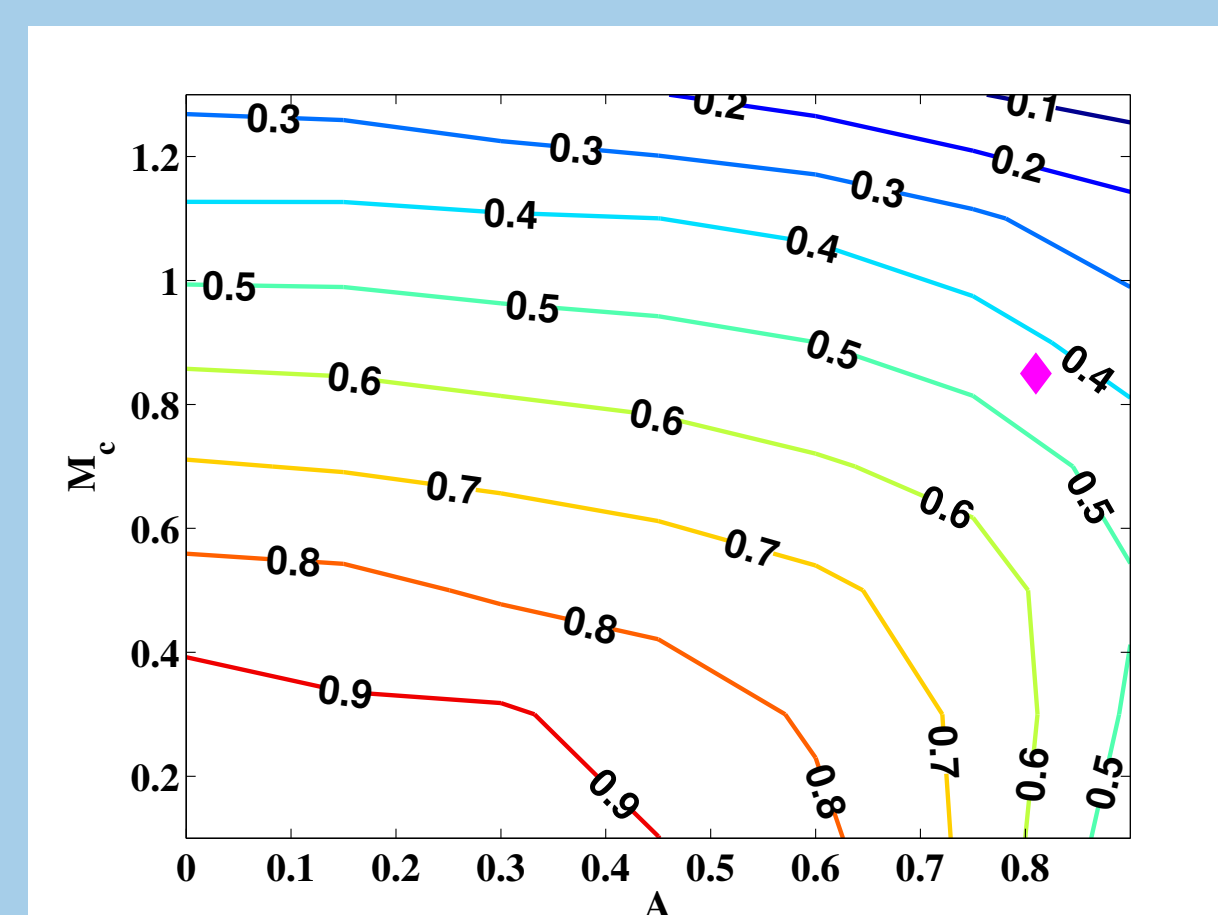


The Kelvin-Helmholtz instability is commonly observed between cloud layers in gas giants, and among the convective processes of the sun.

Left: NASA, Voyager 1 (1979)
Right: NASA/SDO/Astrophysical Journal Letters



- A high convective Mach number inhibits the growth of the KH instability and, in the limit of $M \rightarrow \sqrt{2}$, shuts down growth entirely.

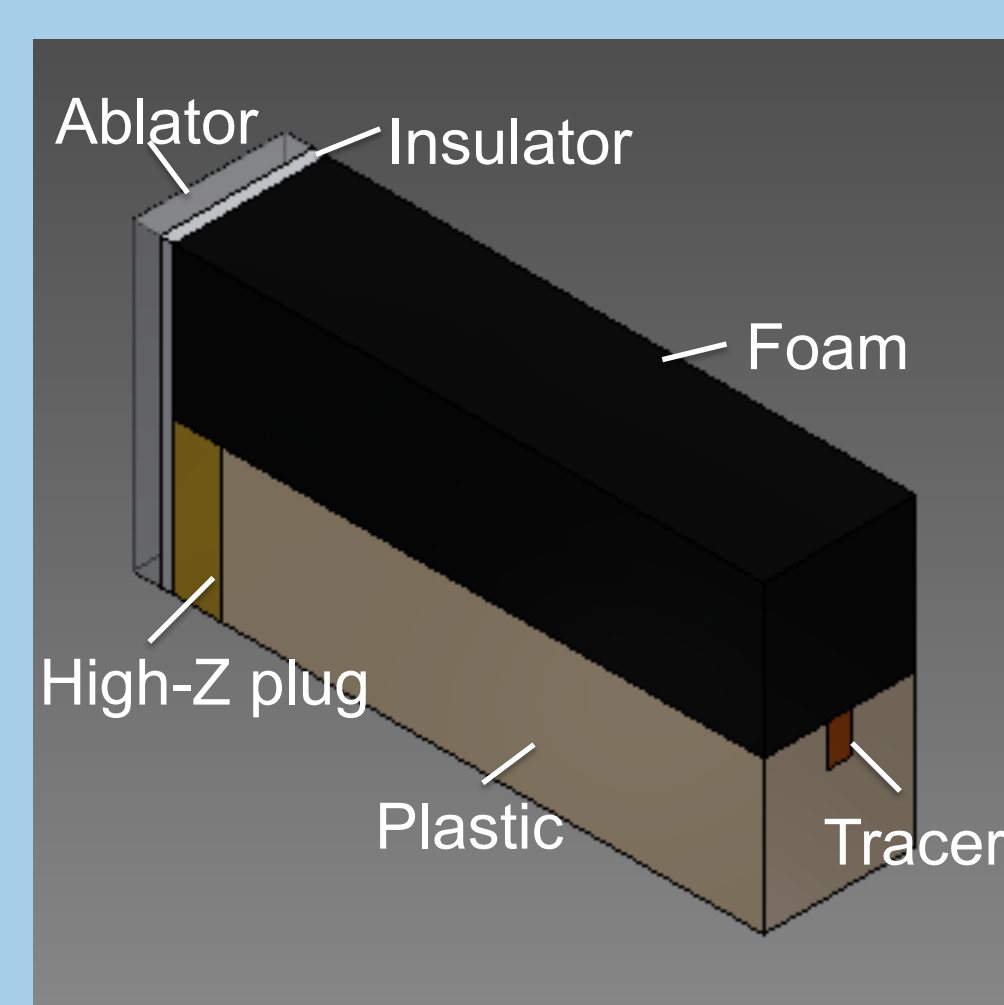
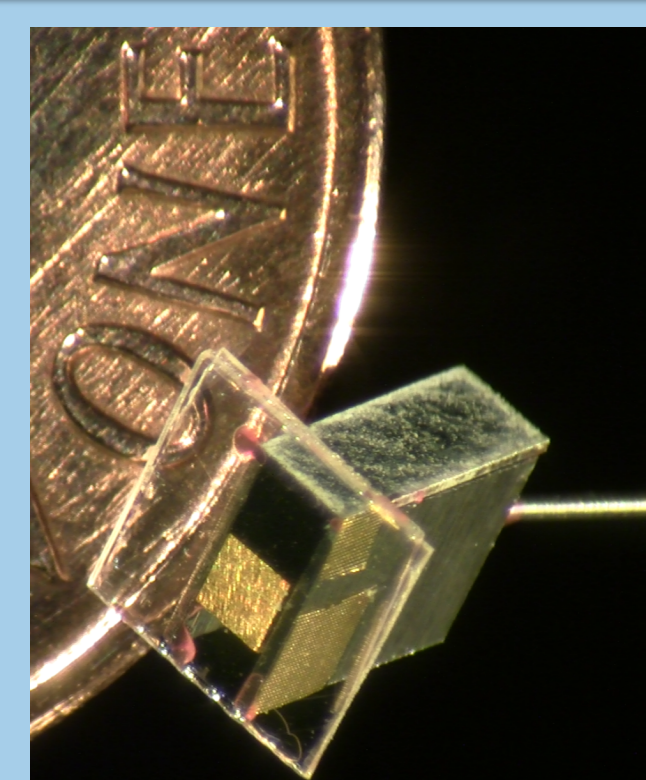


$$\gamma = \frac{k(\Delta u)}{2} \sqrt{1 - A^2} \left(\frac{\sqrt{-1 - M^2} + \sqrt{1 + 4M^2}}{M} \right)$$

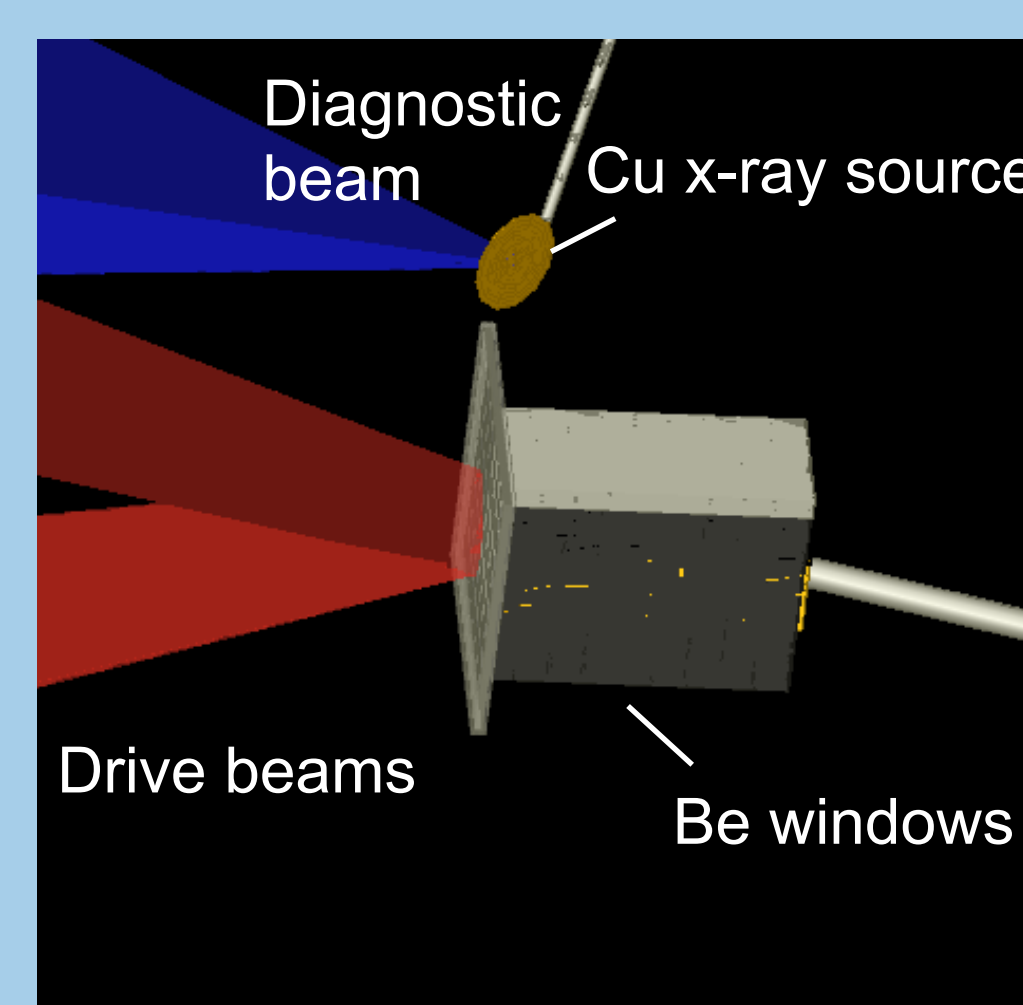
S.R. Choudhury et. al., Journal of Mathematical Analysis and Application. 214, 561-586 (1997).
G. Malamud et al., High Energy Density Physics. 9, 4 (2013)

Shock-wave driven laser hydro experiments

- A 20 to 30 ns pulse drives the system with a steady shockwave, creating shear flow between a low density foam and a high density plastic.



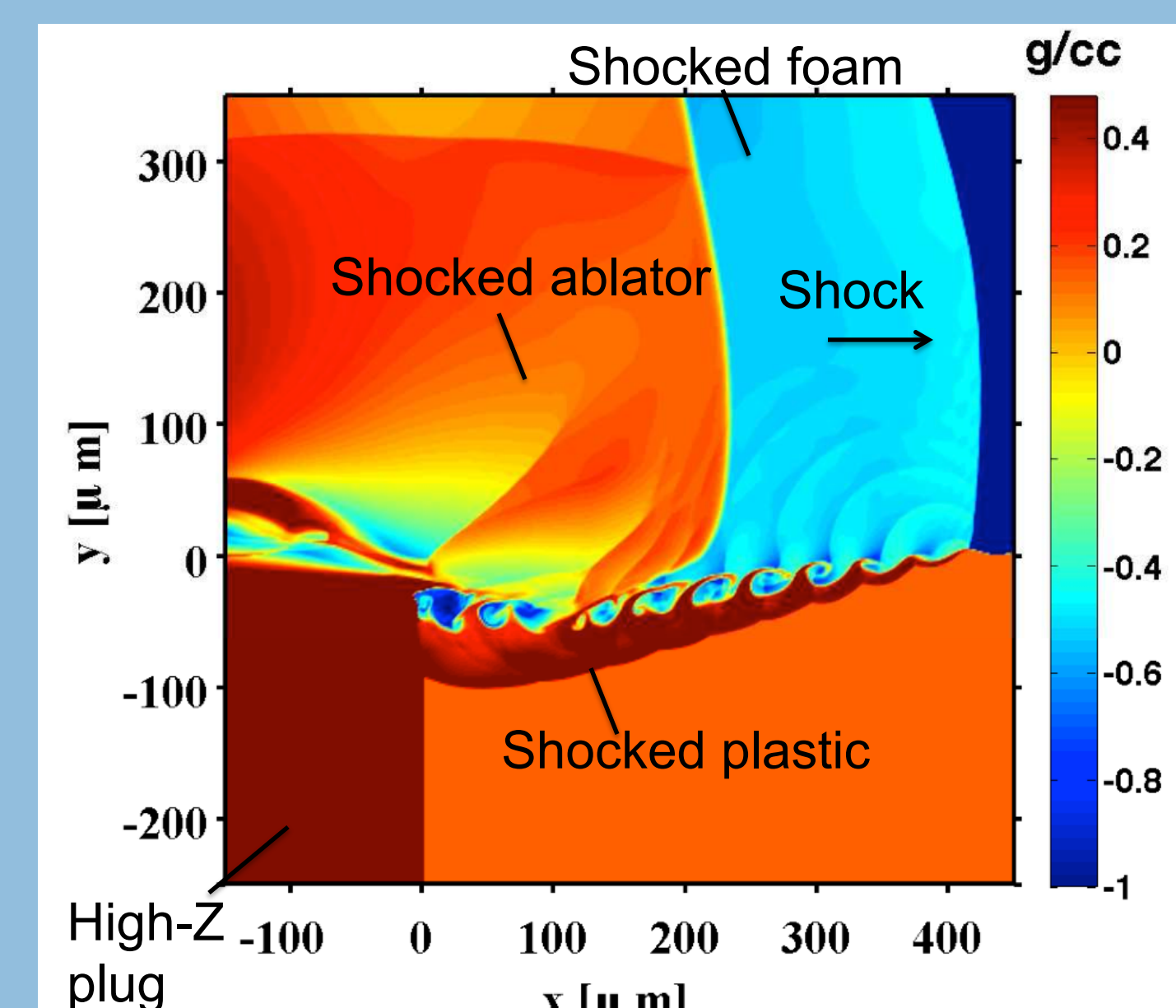
Physics package



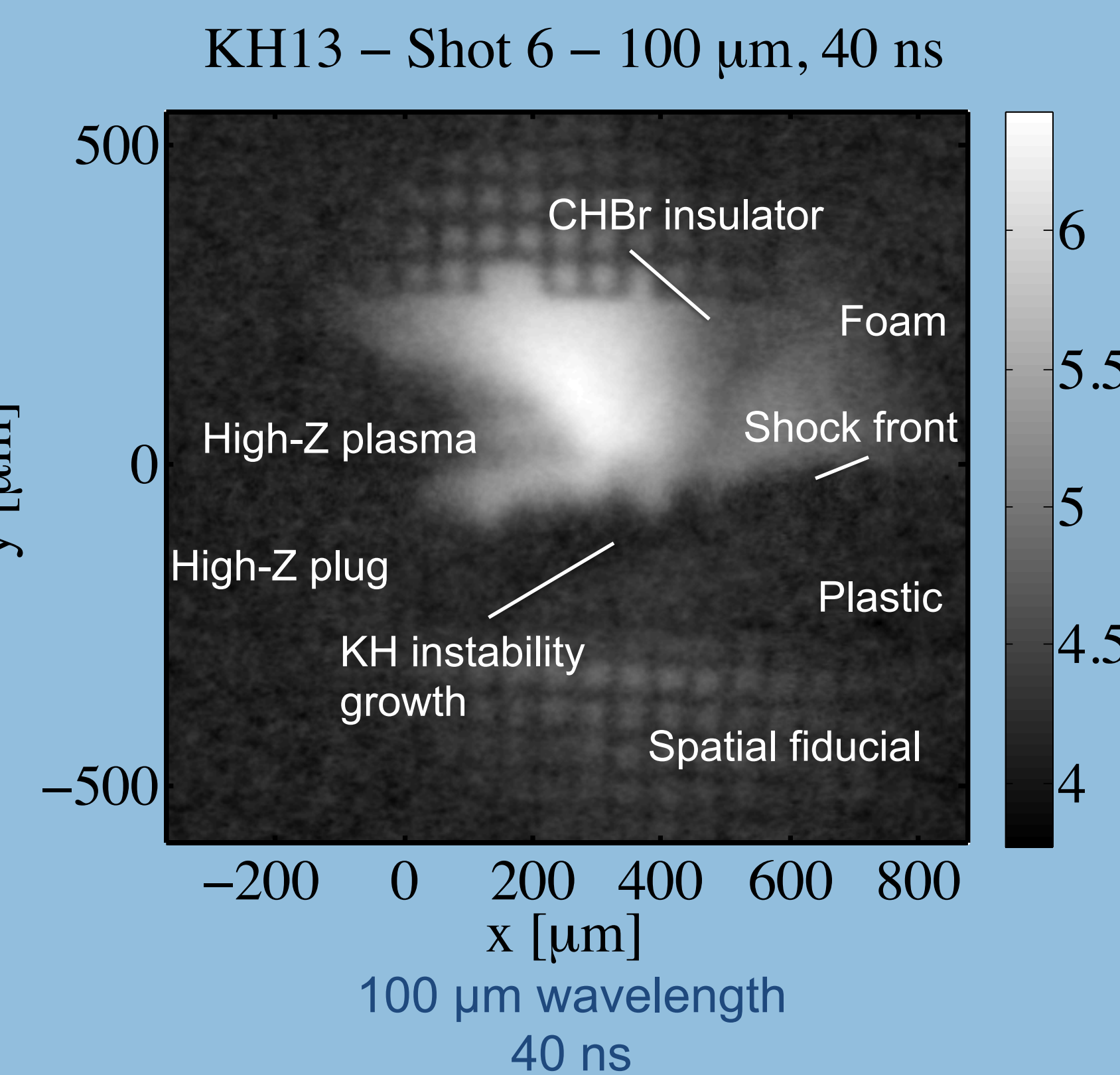
Chamber view

August 2013 results

- In our first campaign, Kelvin-Helmholtz growth may have been limited by a higher than anticipated compression of our foam, resulting in a smaller instability relevant region.



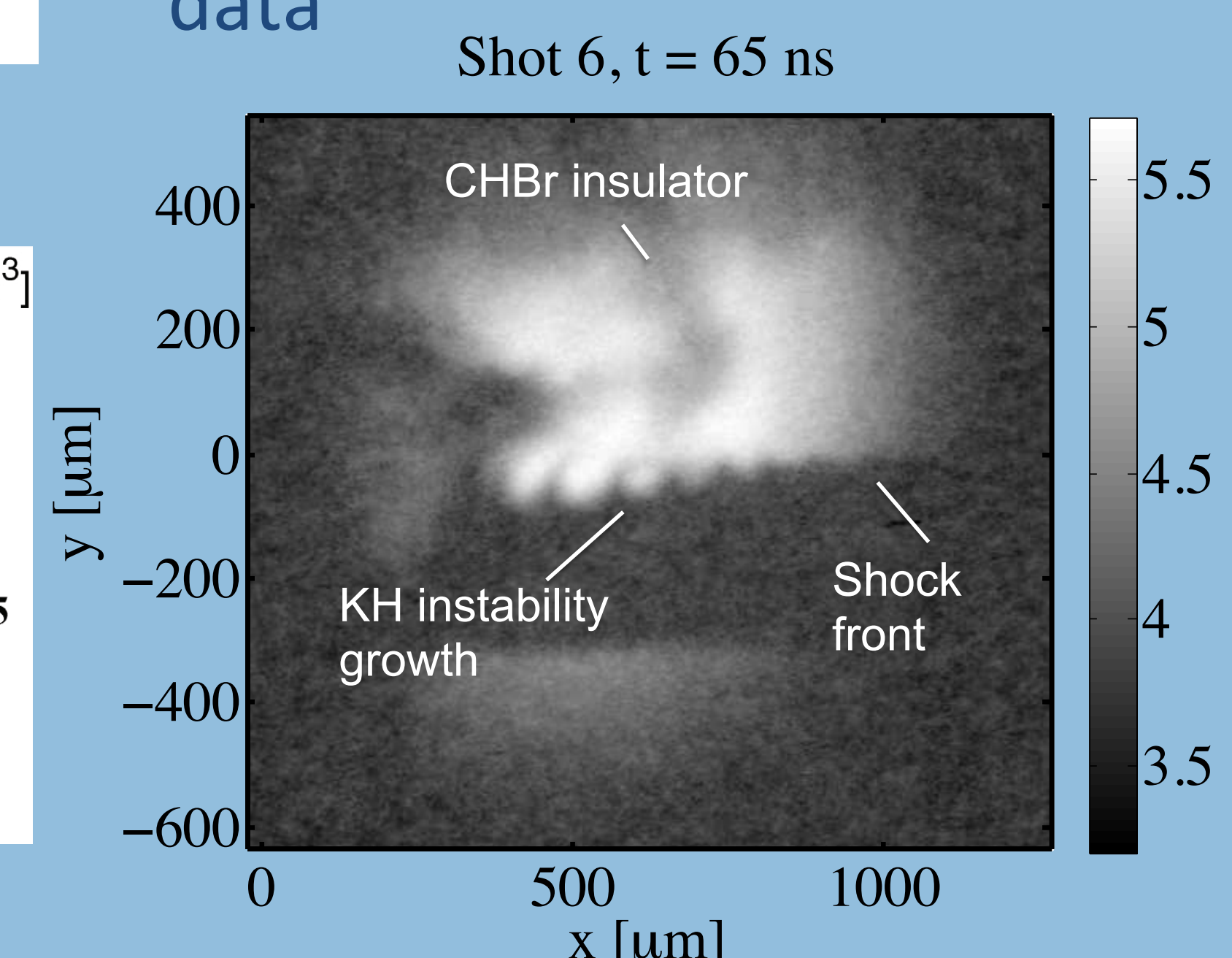
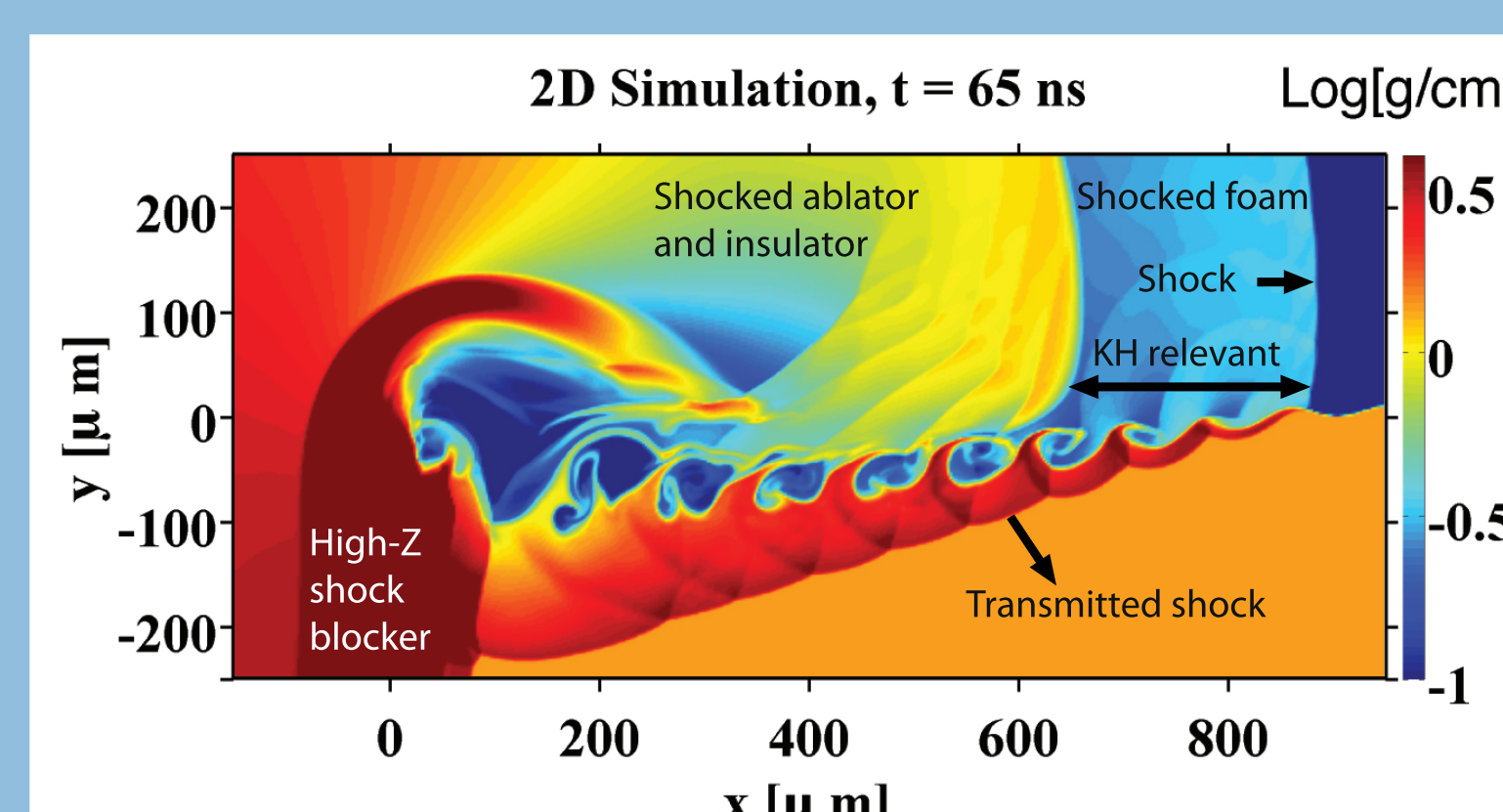
DAFNA 2D simulation
40 ns, log(p)



January 2014 results

- By increasing our drive from 20 ns to 30 ns, we were able to extend the duration of our steady-drive conditions and look later in time.
- At late times, we were able to observe Kelvin-Helmholtz structure.

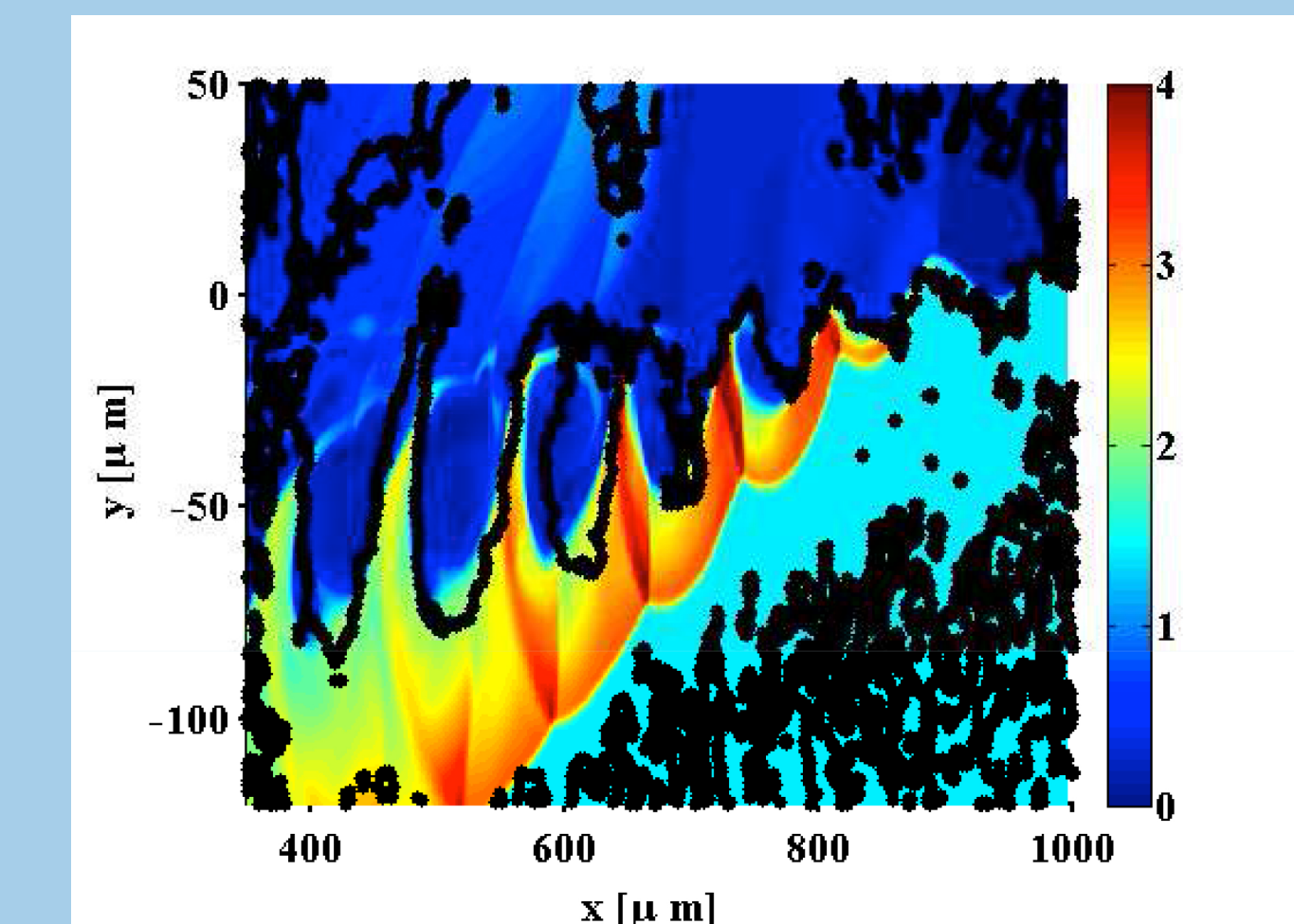
Parameter	Experiment	Simulation
Shock velocity u_s (μm/ns)	28 ± 2	29
Deflection angle (°)	8 ± 2	8.3
Compression	5.0 ± 1.5	4
Shocked foam pressure (Mbar)	0.63 ± 0.08	0.59
Shear velocity Δu (μm/ns)	20 ± 3	20
Rollup spacing λ (μm)	77 ± 6	80



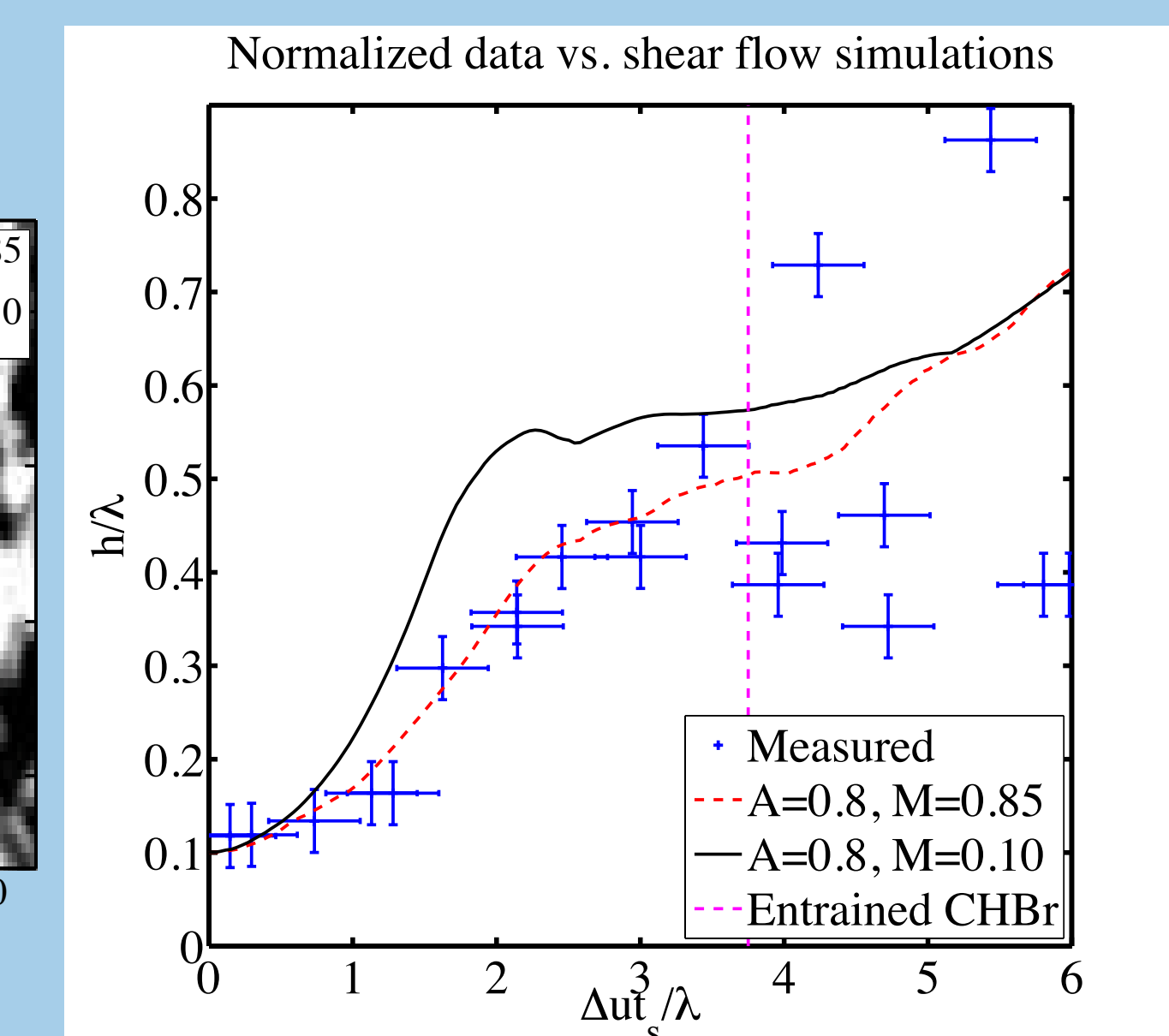
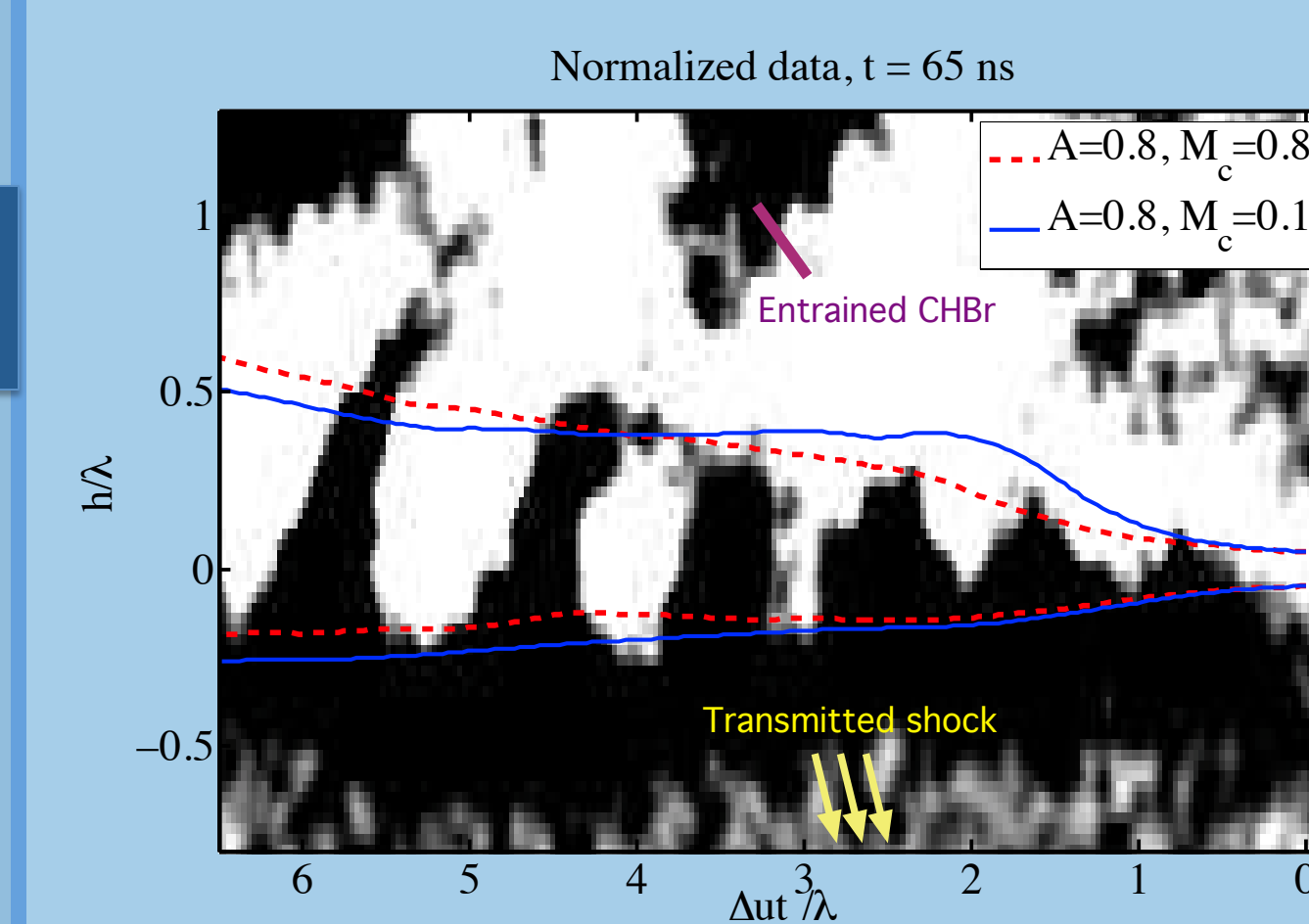
Analysis

- The growth rate of the modulations is in good agreement with simulations, until they are disturbed by entrained material from the ablation surface.

- Top right:** Contour of experimental data plotted over a full 2D simulation.



- Bottom:** Comparison to pure shear flow simulation, at high and low convective Mach numbers.



Conclusions

This data is the first lab observation of single-mode, Kelvin-Helmholtz instability evolution in a supersonic flow.

Funding acknowledgements

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