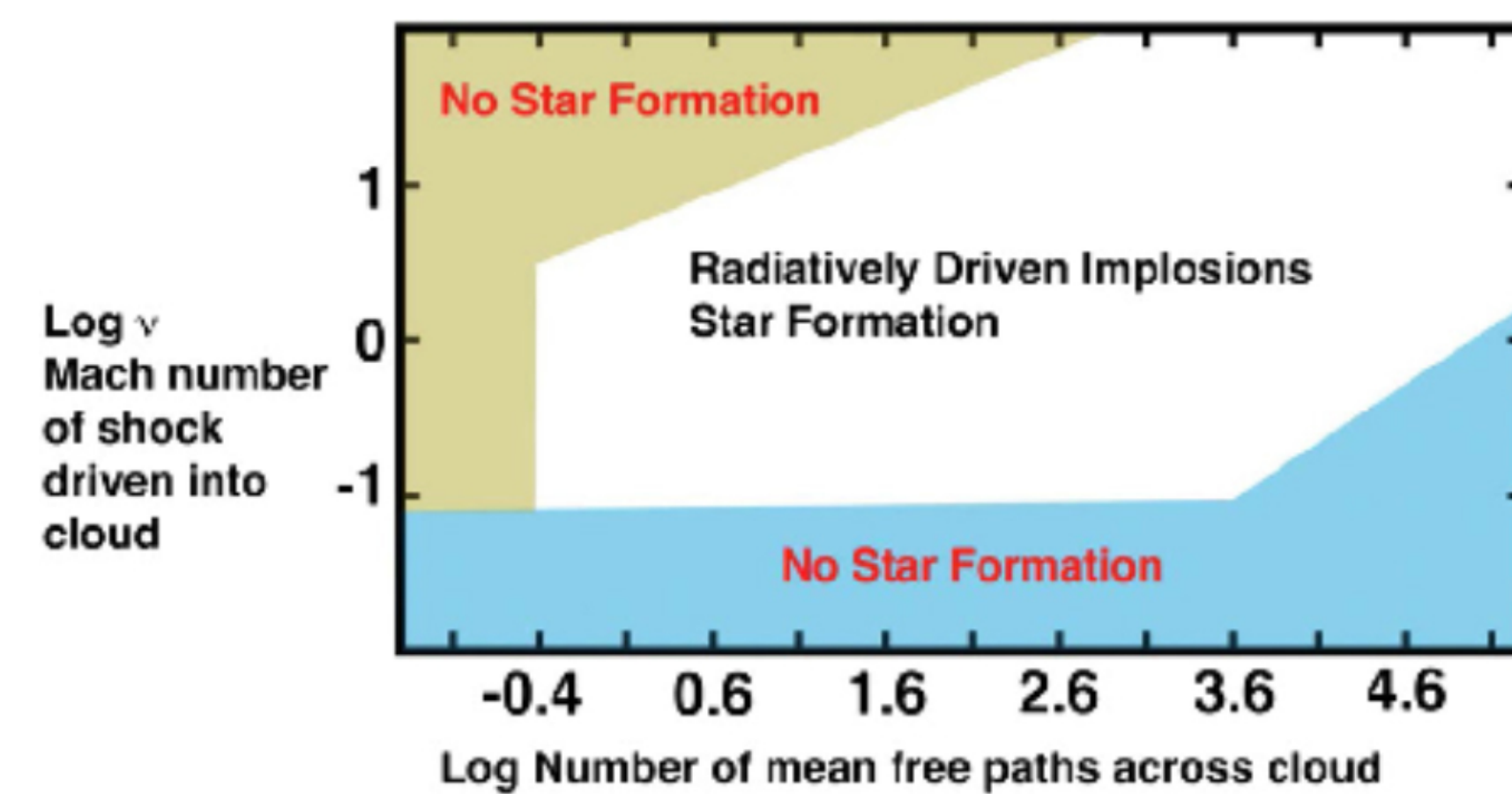


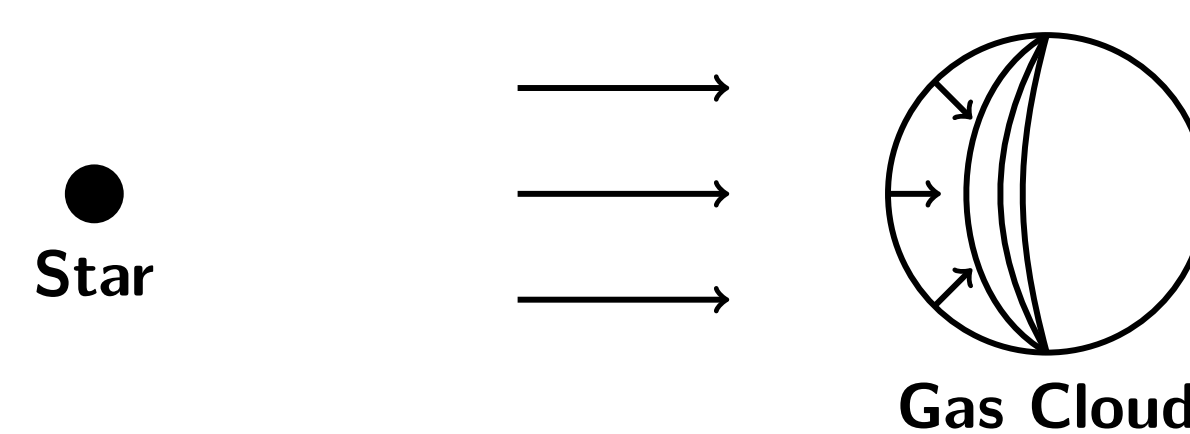
Stellar radiation compresses or explodes local gas clouds



[1] P. A. Keiter et al.: "Design of laboratory experiments to study radiation-driven implosions," High Energy Density Physics, Volume 22, p. 37-40 (2017)

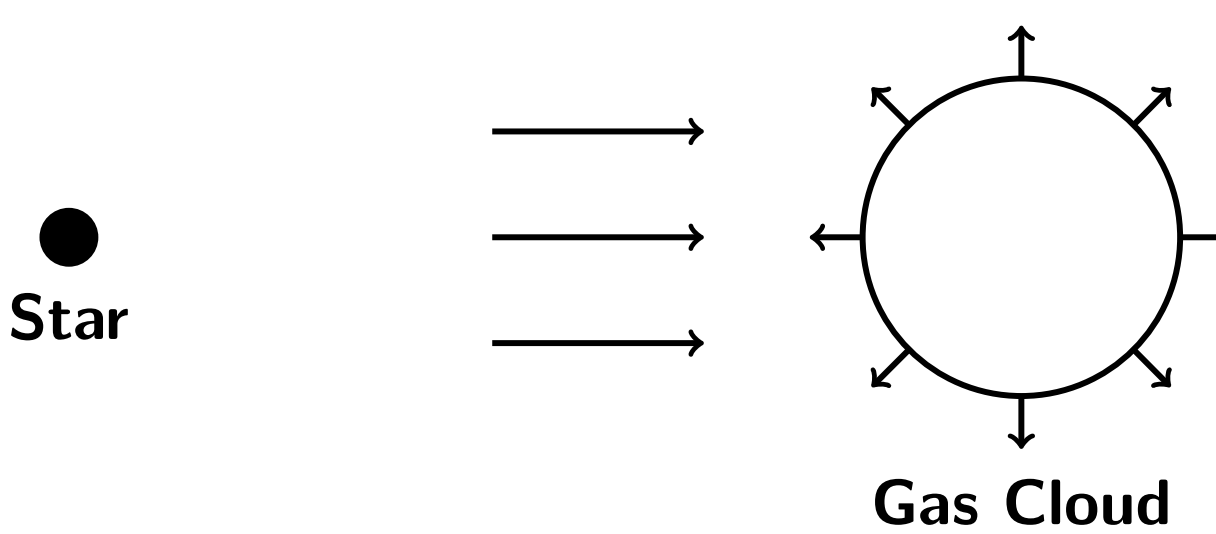
- Interaction dependent on photon energy, cloud thickness and density

Parallel Radiation



- Optically thick: Photons absorbed at one cloud edge
 - Drives asymmetric shock into cloud

Parallel Radiation



- Optically thin: Photons permeate and heat cloud
 - Cloud explodes

Funding acknowledgments

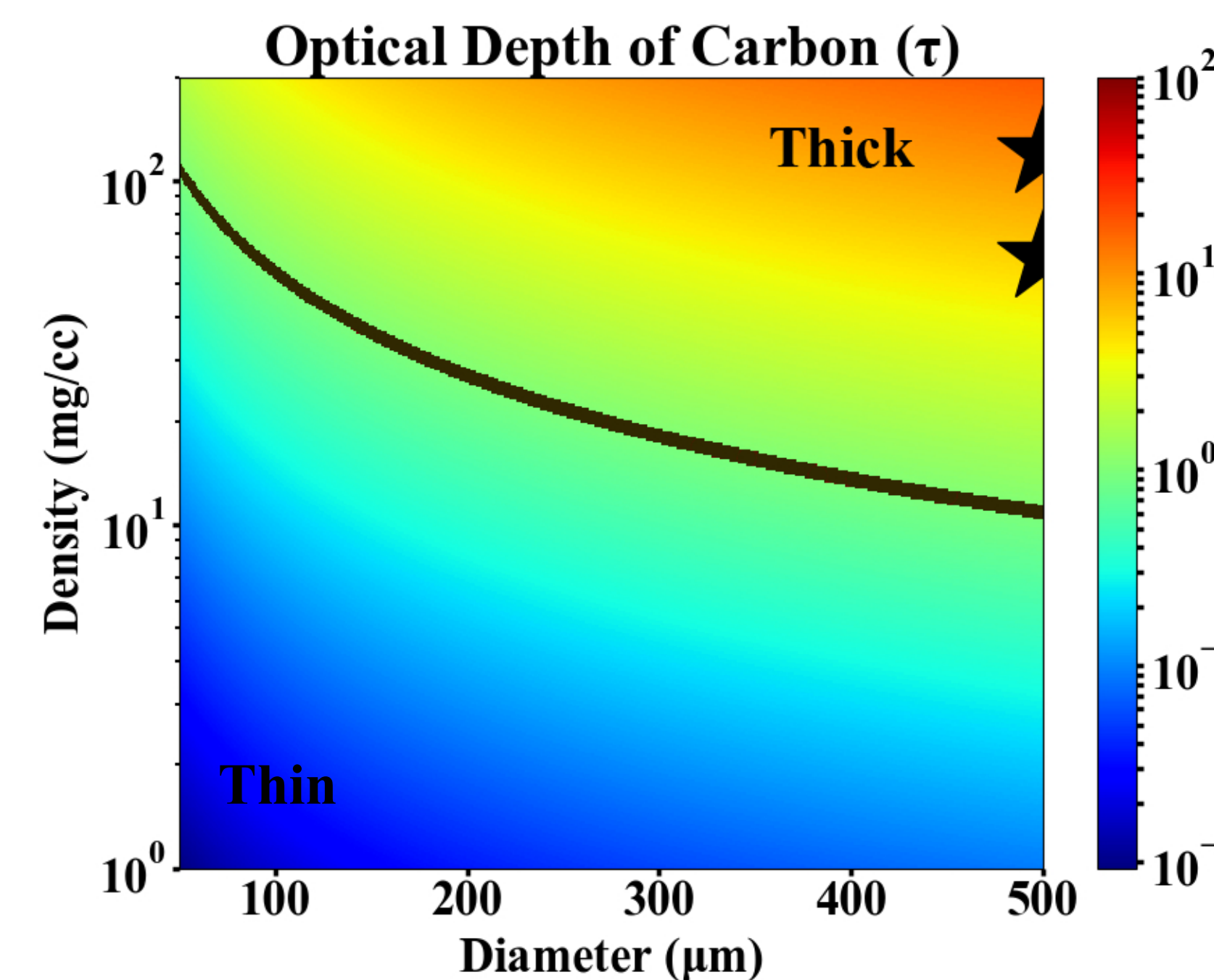
This work is funded by the U.S. DOE, through the NNSA-DS and SC-OFES Joint Program in HEDPLP, grant No. DE-NA0002956, and the NLUF Program, grant No. DE-NA0002719, and through LLE, University of Rochester by the NNSA/OICF under Cooperative Agreement No. DE-NA0001944. This work is funded by the Lawrence Livermore National Laboratory under subcontract B614207.

Additional references

- Drake, R. P. 2010, High-Energy-Density Physics (Springer)
- F. Bertoldi: "The Photoevaporation of Interstellar Clouds. I. Radiation-Driven Implosion," Astrophysical Journal, vol. 346, p. 735-755 (1989)

Foil temperature, foam diameter and density determine the optical depth regime

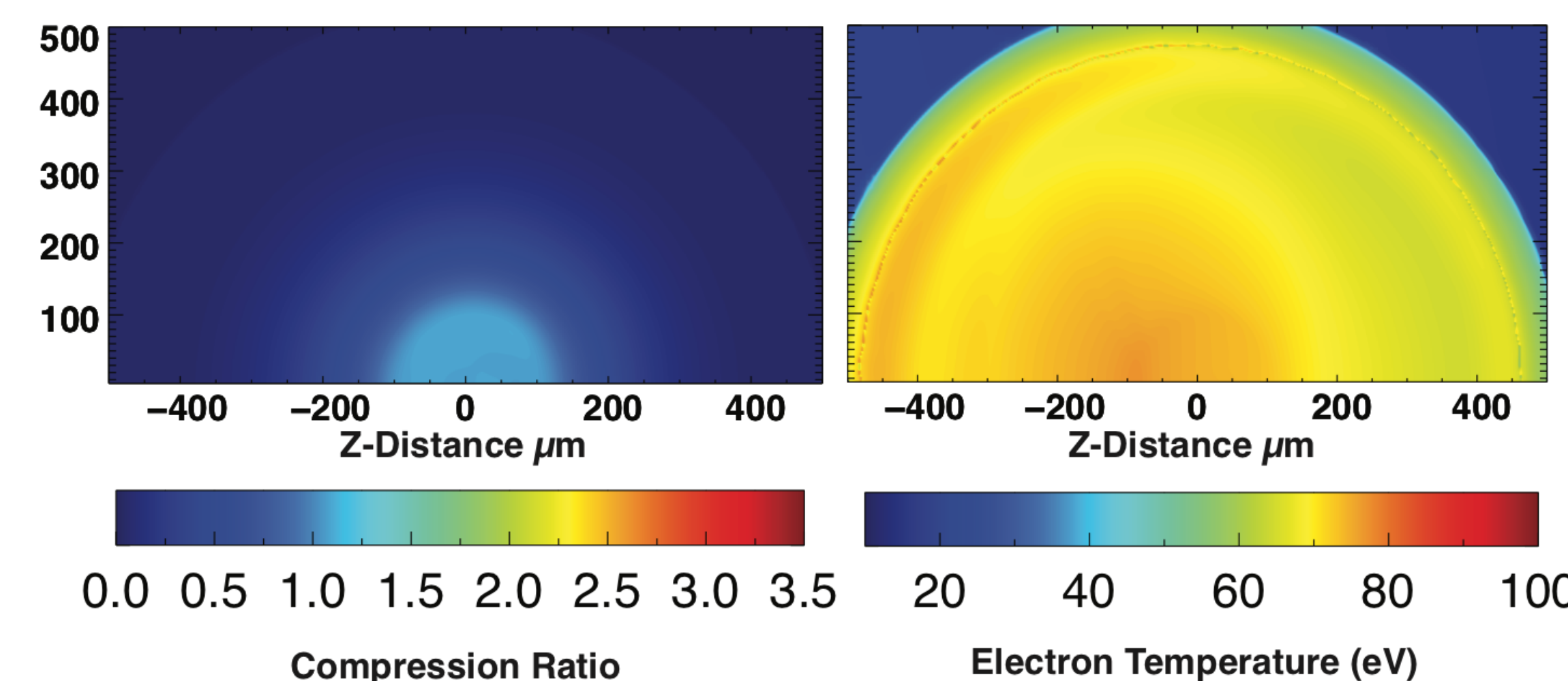
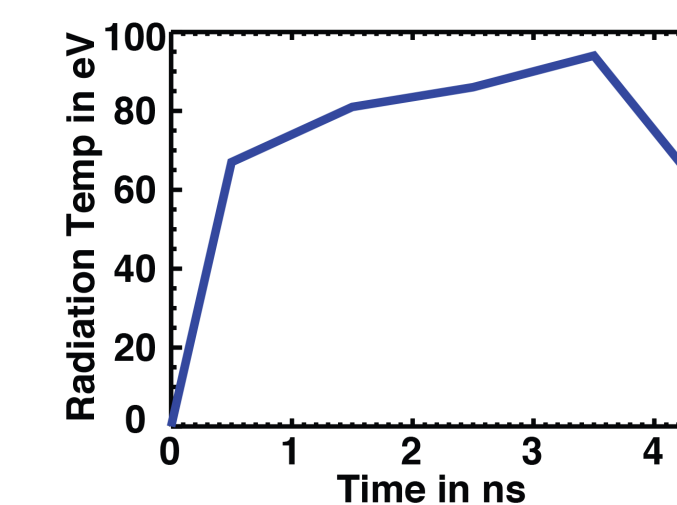
- Carbon foam
- Photon energy ~ 282 eV
- $\sigma \sim 1848 \frac{\text{cm}^2}{\text{g}}$ [2]
- $\lambda_{\text{mfp}} = \frac{1}{\rho\sigma}$
- $\tau \sim \frac{D}{\lambda_{\text{mfp}}} \sim D\rho\sigma$
- $\tau \sim 1$ = black line
- Black stars = experimental parameters



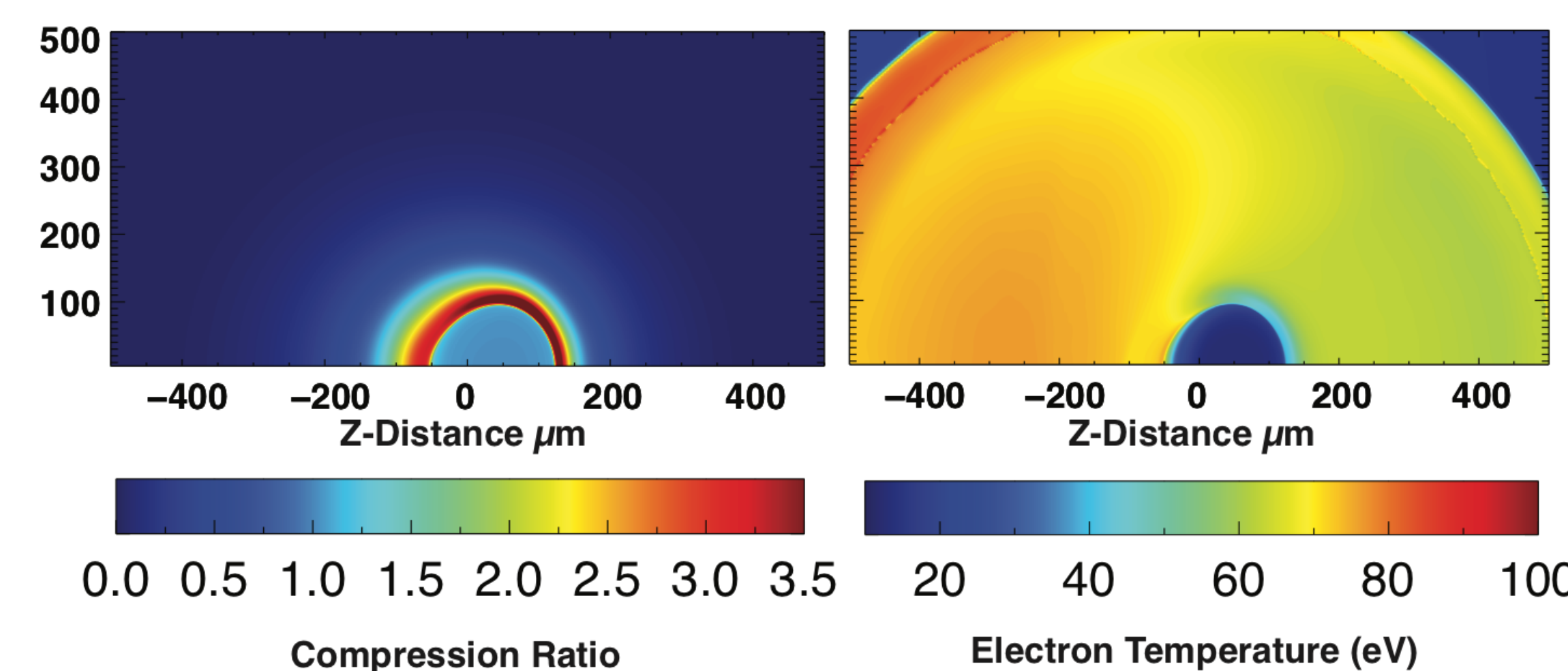
[2] Estimated from http://henke.lbl.gov/optical_constants/filter2.html

2D CRASH radiation hydrodynamic simulations show both compression and explosion limits

- Sphere diameter = 500 μm
- Foil edge to sphere center distance = 1 mm
- Simulated images below are at 2.5 ns
- Radiation source on left side of image



- Optically thin case at 10 mg/cc
- Uniform heating causes uniform explosion

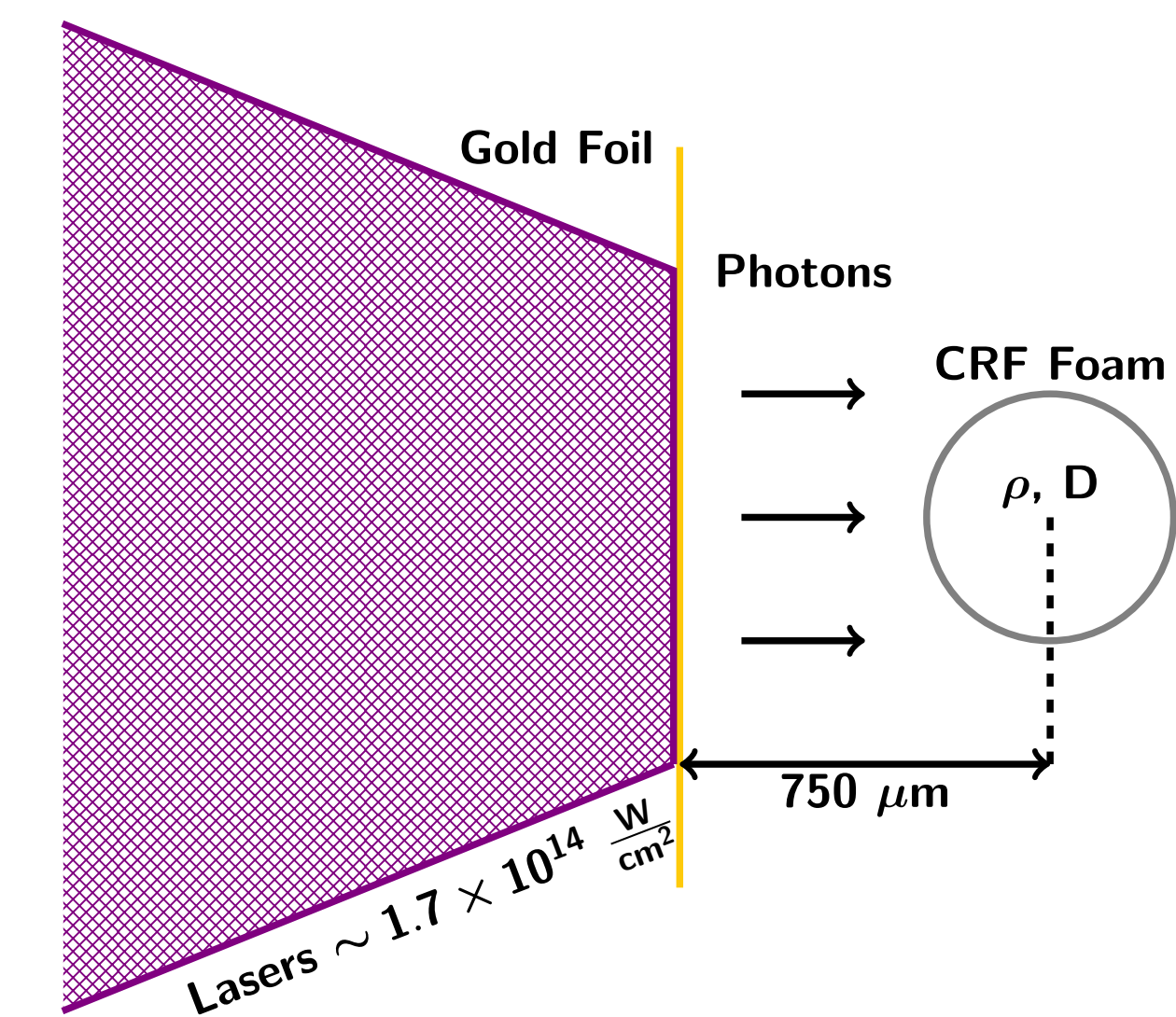


- Optically thick case at 100 mg/cc
- Asymmetric ablation drives an asymmetric compression

[3] B. van der Holst et al; "CRASH: A Block-adaptive-mesh Code for Radiative Shock Hydrodynamics-Implementation and Verification"

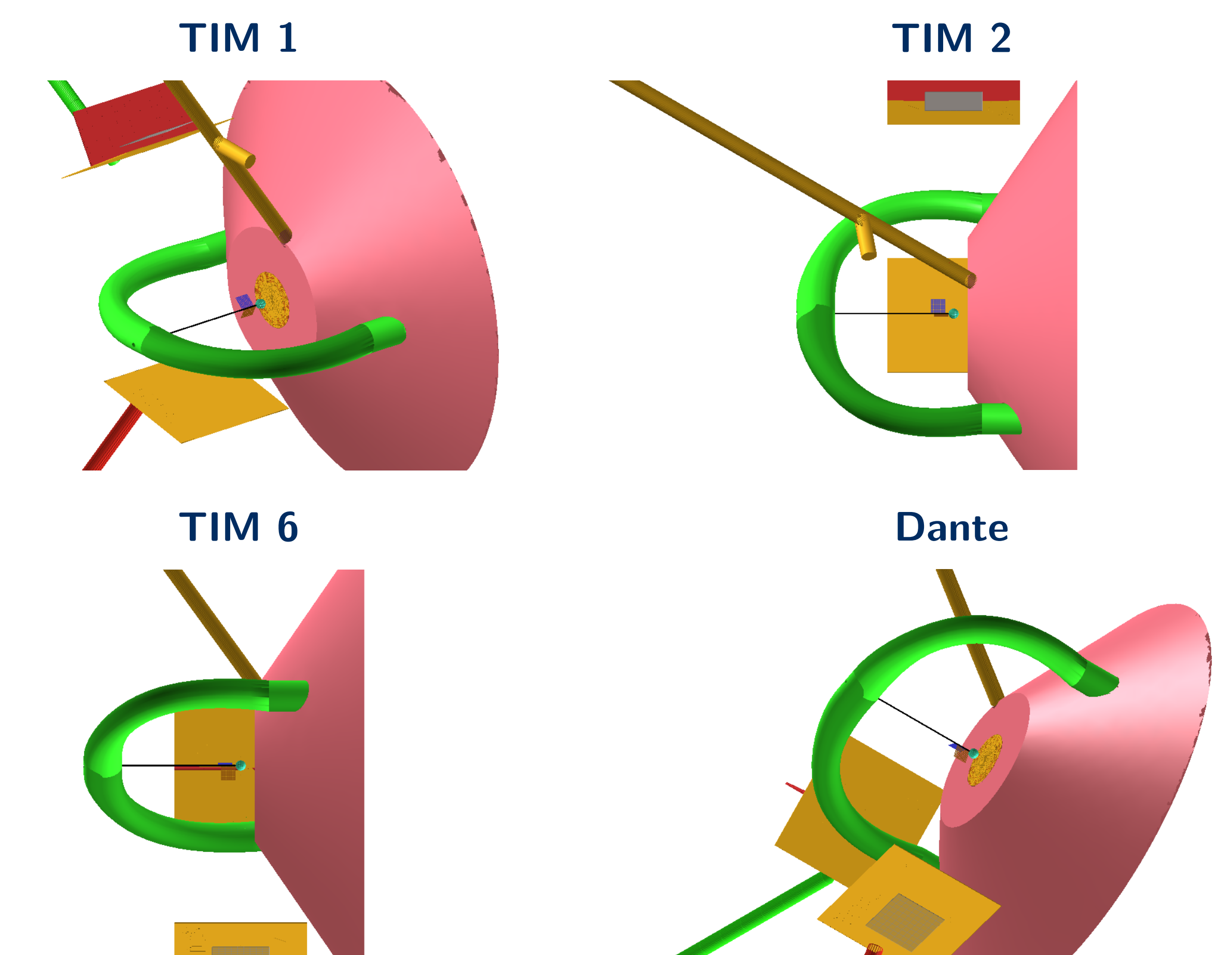
The platform driven by laser irradiated thin gold foil and a variable density foam sphere

- Sphere densities ~ 60 , and $120 \frac{\text{mg}}{\text{cm}^3}$
- Optical depths ~ 5.5 and 11
- Physics drive $\sim 1.7 \times 10^{14} \frac{\text{W}}{\text{cm}^2}$ for 7ns
- Titanium point-projection radiography
 - High densities and late times
- Chlorine point-projection radiography
 - Low densities and early times



[4] J. Davis et al.: "Measurements of laser generated soft X-ray emission from irradiated gold foils," Review of Scientific Instruments 87, 11D609 (2016)

Initial experiments focus on the $\tau \gg 1$ regime



We are developing an experimental platform to explore the range of conditions found when stars irradiate gas clouds

- Local O, B type stars cause gas clumps to explode or compress
 - Gas cloud fate dependent on initial optical depth
- Simulations show the platform can reach optically thin and thick limits by changing foam density
- First experiments started in the optically thick regime for physics and diagnostic simplicity
- Our first optically thick targets had a density ~ 60 , and $120 \frac{\text{mg}}{\text{cc}}$ and a diameter $\sim 500 \mu\text{m}$
- Data is under analysis