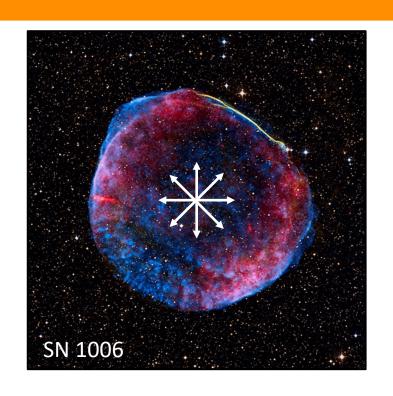
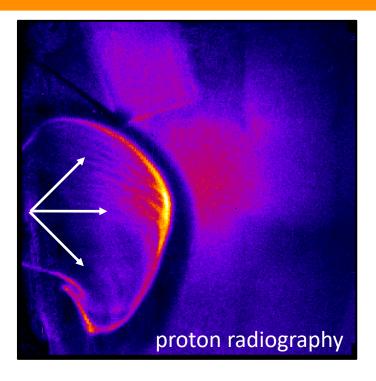
Bringing Cosmic Shock Waves Down to Earth:

Laboratory Studies of Laser-Driven, High-Mach-Number Collisionless Shocks







Derek Schaeffer

Princeton University, Princeton Center for Heliophysics

Early Career Invited Lecturer Series

Michigan Institute for Plasma Science and Engineering

Oct. 21, 2020

Collaborators



- Will Fox (PPPL)
- Russ Follet (LLE)
- Dan Haberberger (LLE)
- Gennady Fiksel (U. Michigan)
- Chikang Li (MIT)
- Jack Matteucci (PU/PPPL)
- Kirill Lezhnin (PU/PPPL)
- Amitava Bhattacharjee (PU/PPPL)
- Daniel Barnak (LLE)
- Suxing Hu (LLE)
- Kai Germaschewski (U. NH)











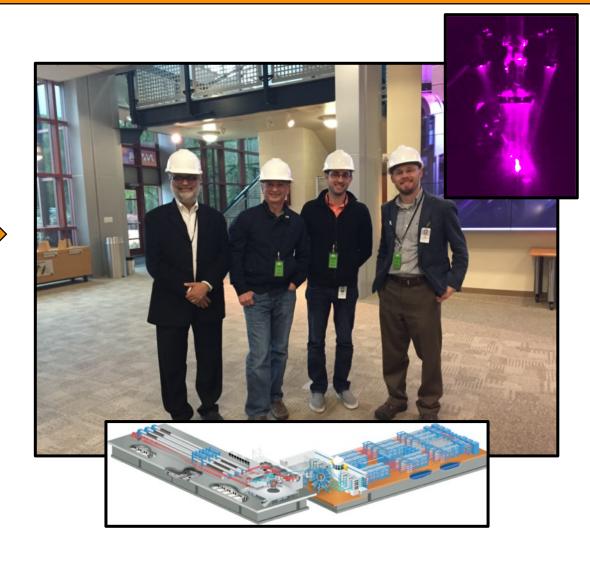


A Brief History





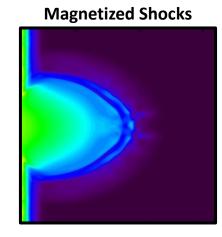




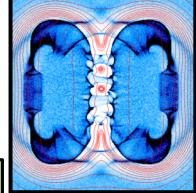
Princeton/PPPL HED Group



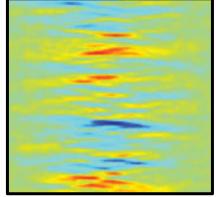
- We study high-energy-density (HED) magnetized plasmas in laser-driven systems
- HED plasmas can often be mapped to vastly different parameter regimes, such as space plasmas, by matching important dimensionless parameters
- Our "laboratory astrophysics" experiments include:
 - Collisionless shocks
 - Magnetic reconnection
 - Ion-scale magnetospheres
 - Biermann-battery magnetic field generation
 - Anomalous/turbulent transport in magnetized plasmas









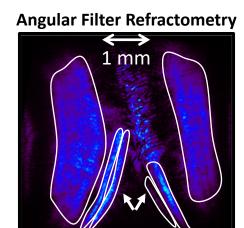


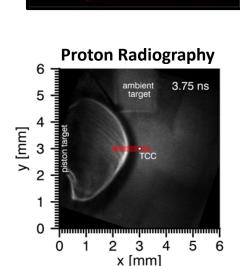
Summary

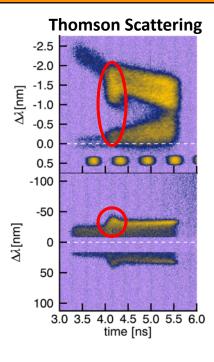


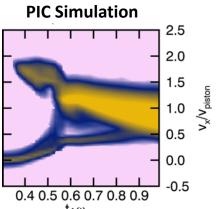


- We have developed a platform for studying laser-driven, high-Mach-number collisionless shocks utilizing advanced diagnostics.
 - First observation of a high-Mach-number magnetized shock in the laboratory
 - First measurements of both ion and electron velocity distributions in a developing magnetized shock
- We have carried out comprehensive particlein-cell simulations that model laser-driven shocks in experimentally-relevant conditions.
 - Observe robust signatures of kinetic-scale shock formation









Outline



- Introduction to magnetized collisionless shocks
- Particle-in-cell simulations of piston-driven shock formation
- Particle velocity distribution measurements in a shock precursor
- Observations of high-Mach-number magnetized shocks
- Conclusions

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- Introduction to magnetized collisionless shocks
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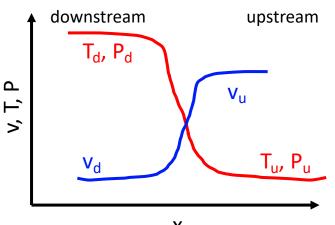
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Shocks Convert Supersonic Ram Pressure to Thermal Pressure









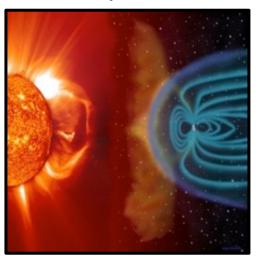
- Incoming supersonic flow is slowed down to subsonic speeds by increasing temperature and pressure
- Characterized by sonic Mach number $M_S = v/c_S$
- Sharp boundary between upstream and downstream regions: shock width on the order of the particle collisional mean free path
- Irreversible process (entropy increases): dissipation provided by collisions
- Energy and momentum conservation yield ratios of upstream to downstream parameters ("jump conditions")

Collisionless Shocks are Prevalent in Many Astrophysical Systems

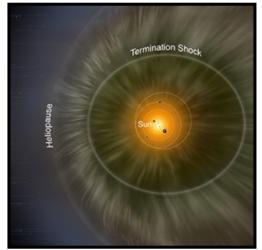




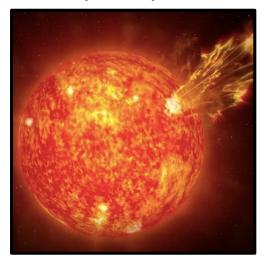
Planetary Bow Shocks



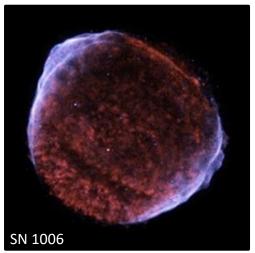
Stellar Bow Shocks



Interplanetary Shocks



Supernova Remnants



- Shocks observed in astrophysical systems with scale lengths orders of magnitude smaller than the collisional mean free path
- Known to be the source of very highenergy particle acceleration, including cosmic rays
- Without collisions, how do they form?

Images: NASA

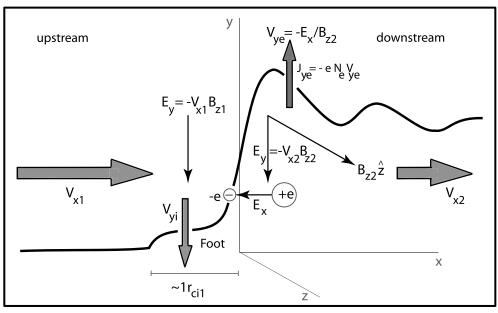
Collisionless Shocks form through Collective Electromagnetic Effects





- Can be categorized as magnetized, electrostatic, or turbulent (e.g. Weibel)
- Dissipation process depends on shock criticality
 - Subcritical ($M_{ms} \lesssim 3$): anomalous resistivity
 - Supercritical: reflected ions
- Classified by magnetic geometry
 - Quasi-perpendicular: $\theta_B > 45^\circ$
 - Quasi-parallel: $\theta_B < 45^{\circ}$
- Shock width of order plasma kinetic scales $(d = c/w_p)$
- Characterized by magnetosonic Mach number $M_{ms}=v/v_{ms},\,v_{ms}^2=v_A^2+c_s^2$

Schematic of a Supercritical Perpendicular Magnetized Collisionless Shock

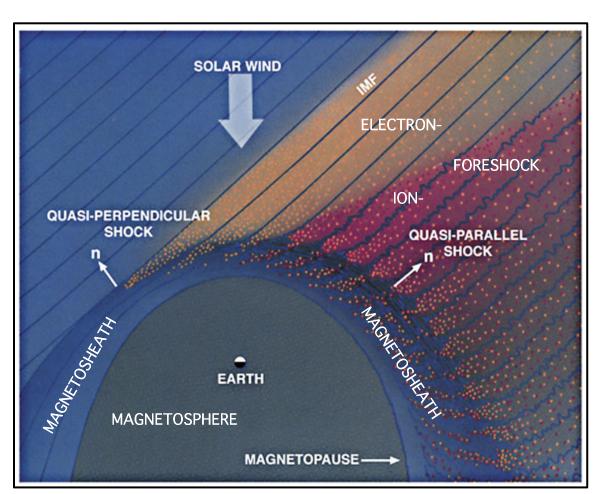


[Balogh & Treumann 2013]

Earth's Bow Shock is a Natural Laboratory for Studying Magnetized Collisionless Shocks







[Tsurutani & Stone 1985]

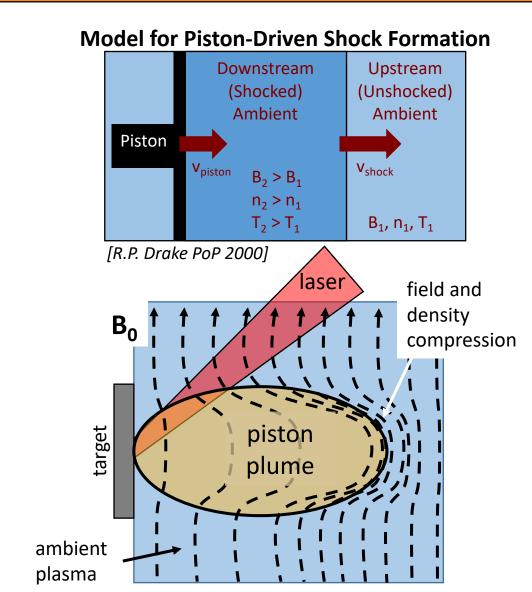
- Very successful satellite program has yielded a wealth of information on shocks
- But spacecraft studies are limited:
 - ~1D trajectories
 - Variable and uncontrolled plasma parameters
 - Focused on small-scale structure
- Telescope observations are limited to mostly large-scale structure
- Many questions remain unanswered
 - How is energy partitioned between electrons and ions across a shock?
 - How are particles injected into shock acceleration mechanism?
 - What are the characteristic scales of shock formation and reformation?
 - What is the role of turbulence and reconnection in high-Mach number shocks?

Laboratory Platforms Allow Detailed Studies of Collisionless Shocks using Laser Plasmas





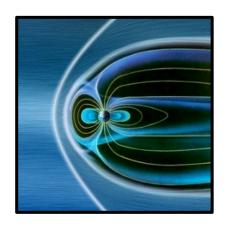
- Laboratory collisionless shocks can be generated by driving a supersonic piston plasma through a magnetized ambient plasma
 - Controlled and reproducible parameters
 - Wide range of Mach numbers $(M_{ms} < 40)$
 - 2D and 3D datasets
 - Flexible magnetic geometry
 - Velocity distribution measurements, which will eventually allow direct comparisons between space and laboratory data

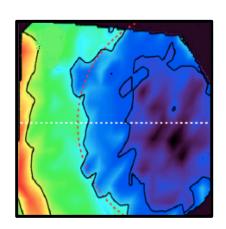


Dimensionless Parameters are Similar in Space and Laboratory Plasmas



Plasma Parameter	Earth's Bow Shock	Laboratory (LAPD)	Laboratory (HED)
Magnetic Field B ₀	5x10 ⁻⁵ G	300 G	4x10 ⁴ G
Initial Piston Expansion Speed v ₀	400 km/s	250 km/s	700 km/s
Upstream Ion Density n ₀	5 cm ⁻³	5x10 ¹² cm ⁻³	6x10 ¹⁷ cm ⁻³
Upstream Ion Inertial Length c/ω_{pi}	100 km	20 cm	240 μm
Shocked Ion Gyroperiod ω_{ci}^{-1}	0.5 s	200 ns	1.5 ns
Shocked Ion Directed Gyroradius $oldsymbol{ ho}_{a}$	200 km	5 cm	1 mm
System Size D ₀	100-1000 km	60 cm	1 cm
Dimensionless Parameter			
Alfvénic Mach Number M _A	8	2	17
Ion Collisional Length Scale λ_{mfp}/D_0	105	10 ²	10
Density (B) Compression n/n_0 (B/B ₀)	2-4	2	4
Shock Ramp Width ∆ x/d _i	~1	~1	~1





Outline



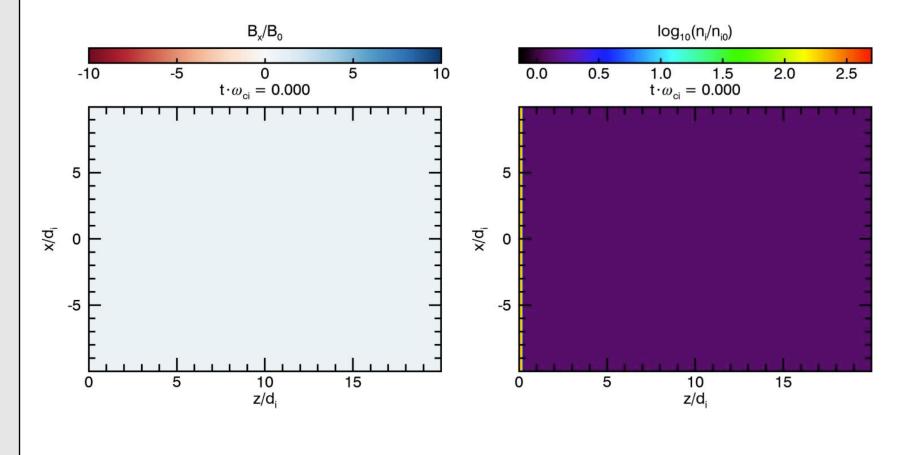
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Experimental System Simulated with Particle-in-Cell Code PSC





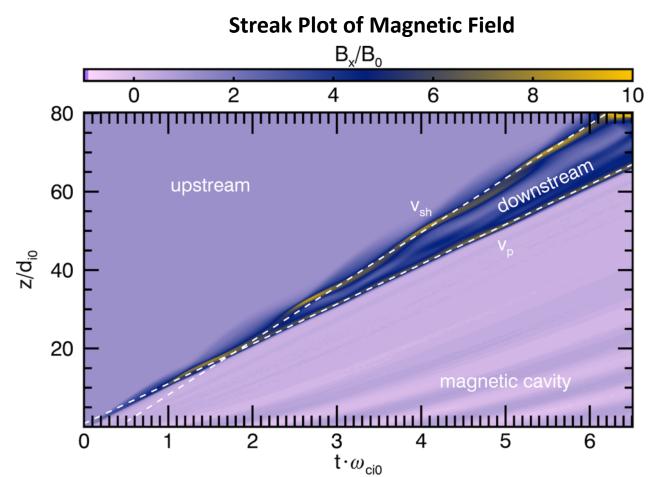
- 2D explicit PIC with Coulomb collision operator
- Heating operator mimics laser ablation and generates supersonic piston plasma
 - Validated against experiments and radhydro codes
- Piston expands through uniform magnetized ambient (upstream) plasma
- Can simulate multispecies plasmas



[Germaschewski+ JCP 2016] [Fox+ PoP 2018]

Simulations Demonstrate Piston-Driven Collisionless Shocks

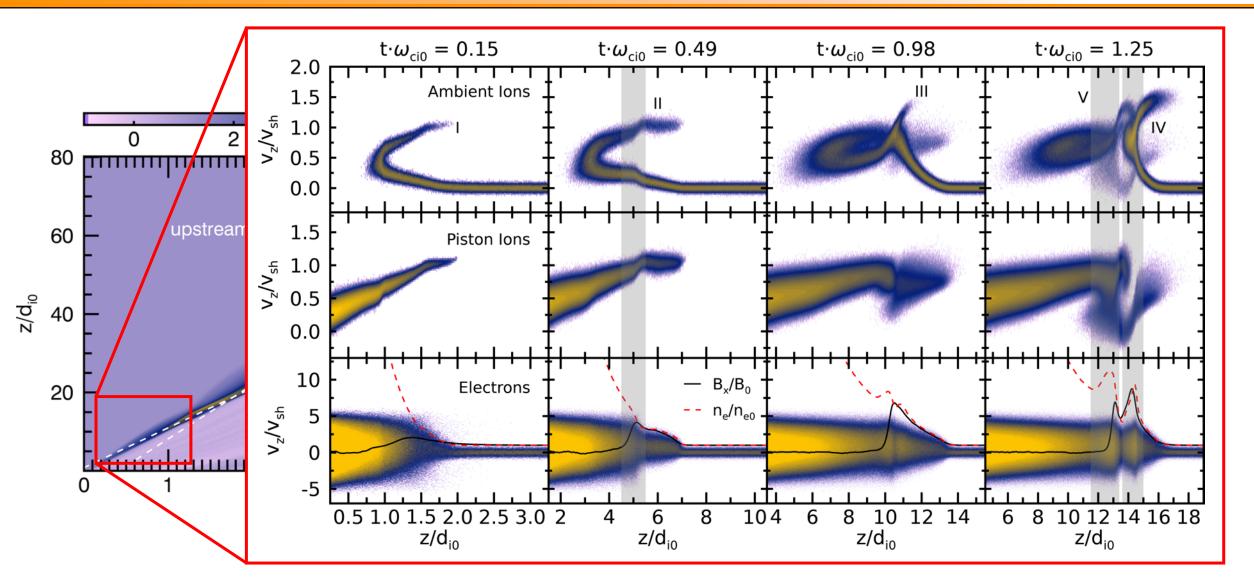




- H⁺¹ ions for piston and ambient plasma
- Perpendicular magnetic geometry

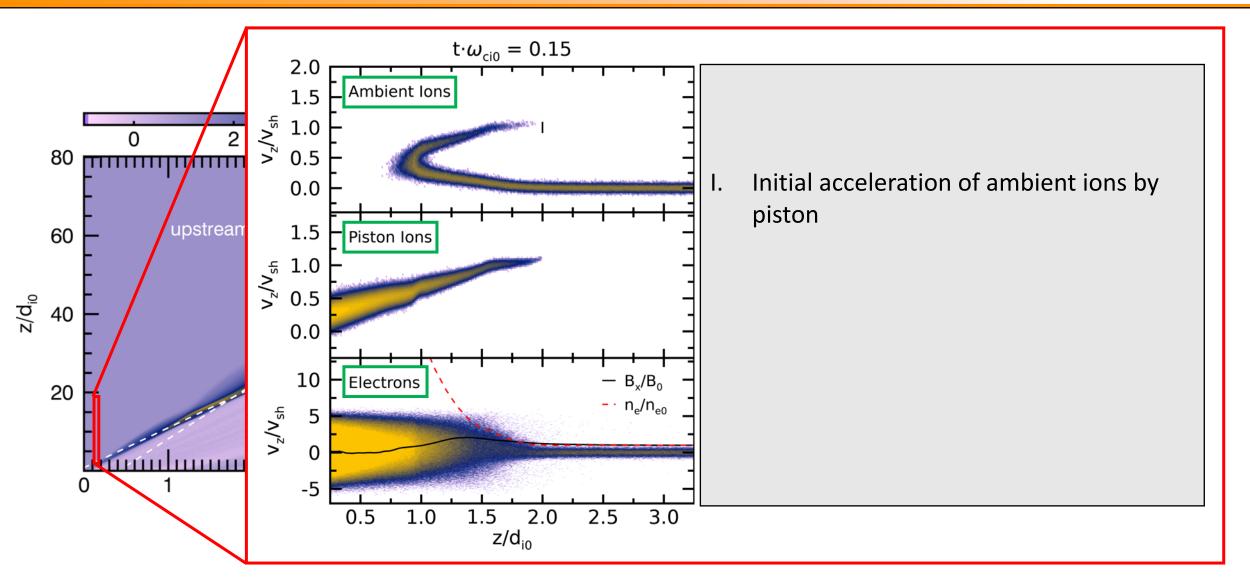






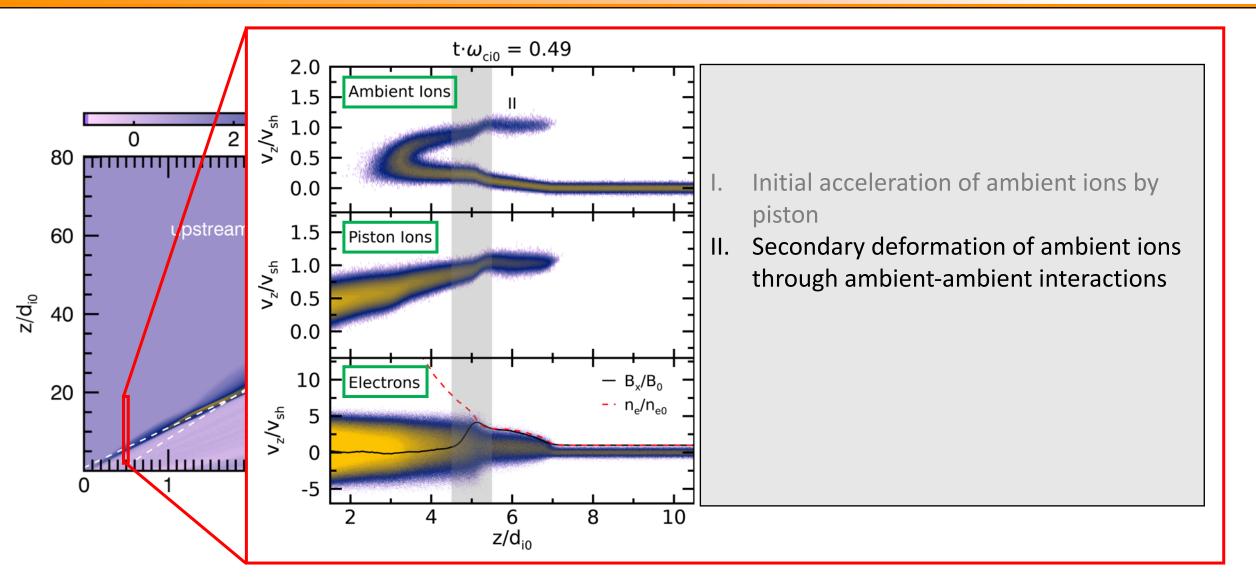






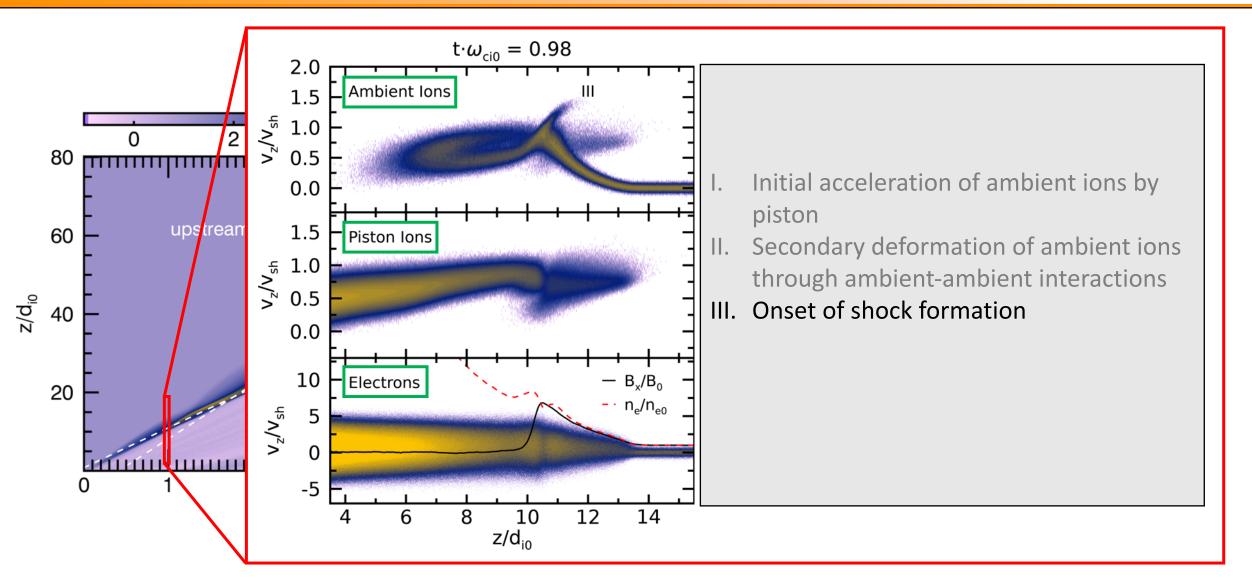






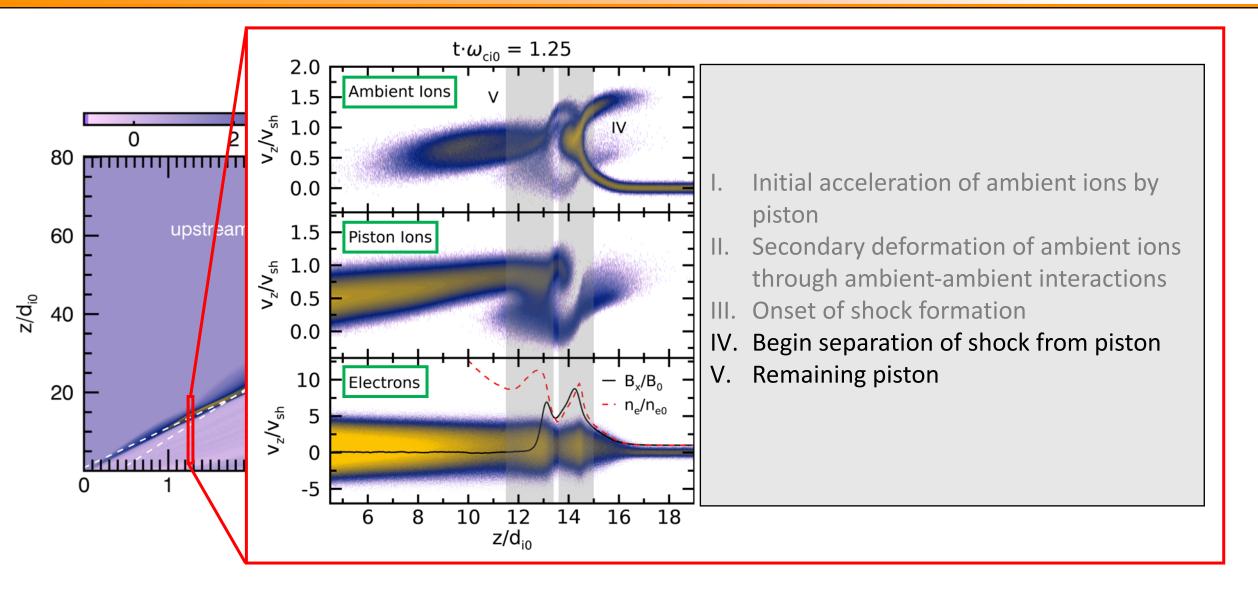








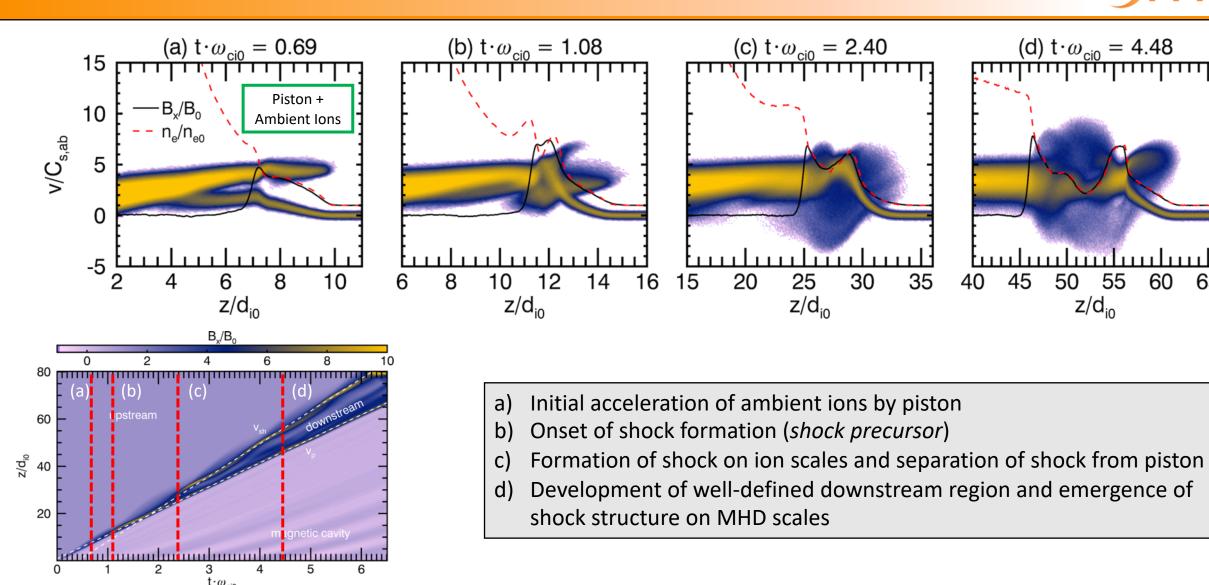




Piston-Driven Magnetized Shocks Form in Three Stages







Piston-Driven Shock Formation is a Complex Process

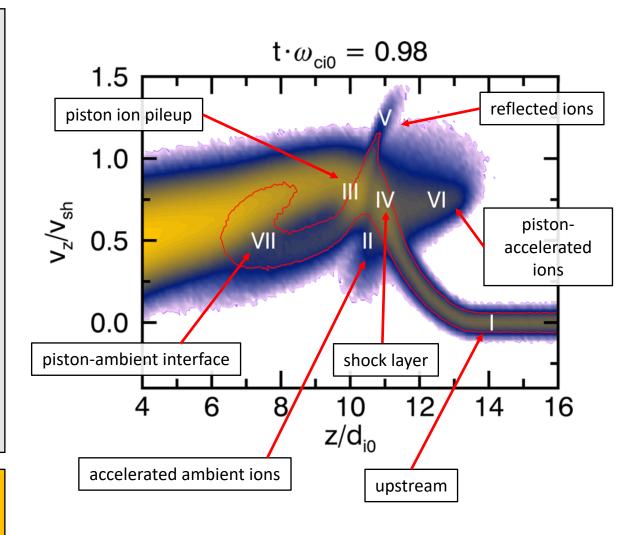




Signatures of shock formation include

- Onset of shock formation ($\sim 1\omega_{ci}^{-1}$)
 - Deformation of piston and upstream flows
 - Strong density and magnetic compressions
 - Upstream ion reflection from compressed magnetic fields
- Piston-shock separation ($\sim 1-2.5\omega_{ci}^{-1}$)
 - Double-bump structure in density profiles
- Development of a downstream region ($\sim 2.5 5\omega_{ci}^{-1}$)
 - Consistent with RH jump conditions
- Important to distinguish piston-dominated and shockdriven processes
 - Magnetic and density compressions necessary but not sufficient conditions

Key experimentally-relevant observables can be extracted from simulations



Outline

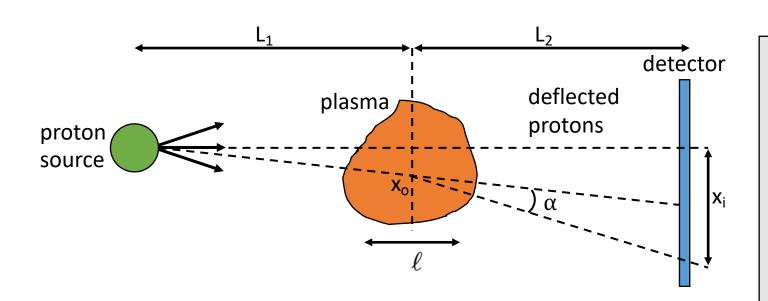


A brief interlude to discuss diagnostics

Proton Radiography Measures Path-Integrated Magnetic Field

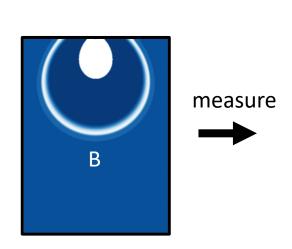


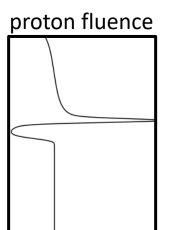


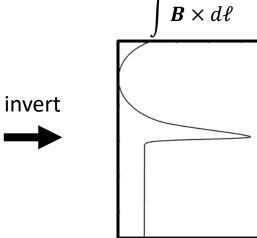


- Protons generated by intense laser interaction with Cu foil (TNSA) or implosion of DHe3 capsule
- Magnetic fields in plasma deflect protons (electric fields negligible)
- Deflected protons collected at detector (image plate or CR-39)
- Signal can be inverted to estimate pathintegrated magnetic field

$$\alpha = \frac{e}{m_p v_p} \int \mathbf{B} \times d\ell$$
$$x_i = x_o \left(1 + \frac{L_2}{L_1} \right) + L_2 \alpha$$







Optical Thomson Scattering Measures Plasma Parameters



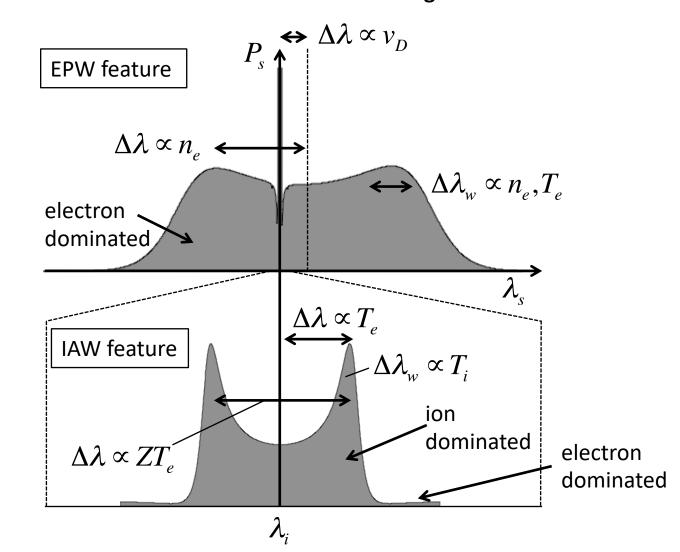


 Collective TS measures scattered light from electron plasma waves (EPW) and ion acoustic waves (IAW)

$$S(\mathbf{k},\omega) = \frac{2\pi}{k} \left| 1 - \frac{X_e}{\epsilon} \right|^2 f_{e,0}(\omega/k) + \frac{2\pi Z}{k} \left| \frac{X_e}{\epsilon} \right|^2 f_{i,0}(\omega/k)$$

- From EPW features, one can extract electron density and temperature
- From IAW features, one can extract flow speed, electron temperature, and ion temperature
- Parameters obtained by iteratively fitting an analytic scattered spectrum to the data

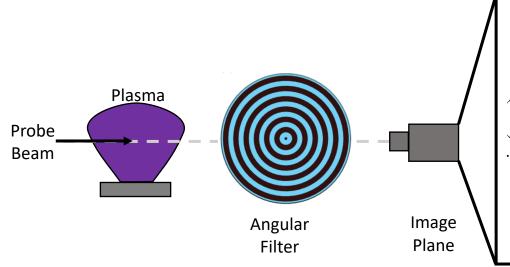
Collective Thomson Scattered Signal



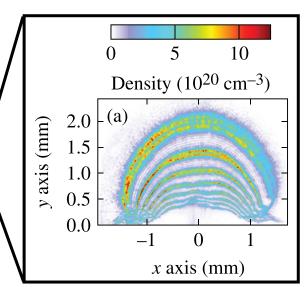
Angular Filter Refractometry Measures Path-integrated Density Gradient Contours



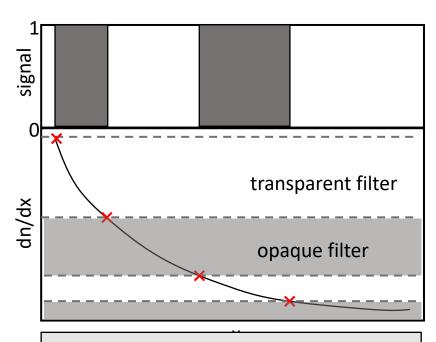




- Light from probe laser is collected an passed through angular filter placed in Fourier plane
- The result is a discrete set of bands corresponding to specific plasma refraction angles
- The angles are proportional to the pathintegrated plasma density gradient



$$\theta_{\alpha} = \frac{1}{2n_{cr}} \int_{-\infty}^{\infty} \frac{\partial n_e}{\partial \alpha} dz$$



- Narrow bands = large change in density gradient
- Broad bands = small change in density gradient

[Haberberger+ PoP 2014]

Outline



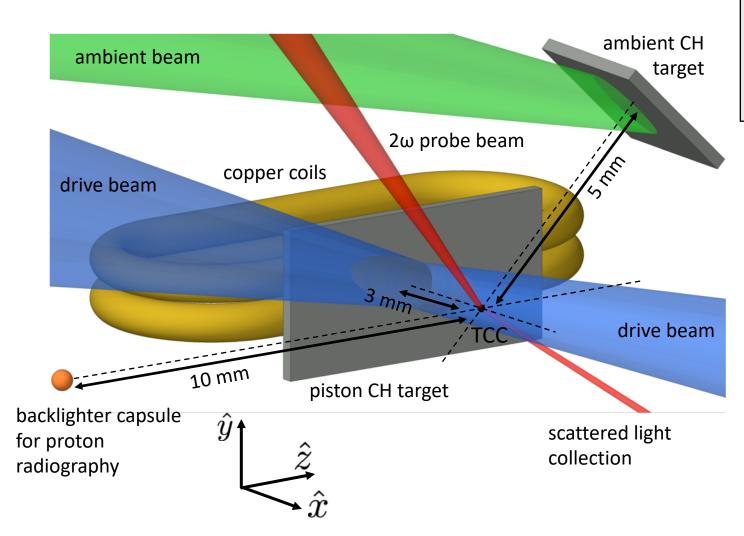
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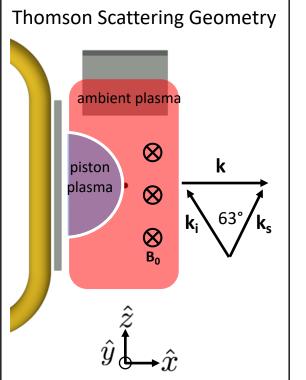
Experimental Setup for Collisionless Shocks on OMEGA 60







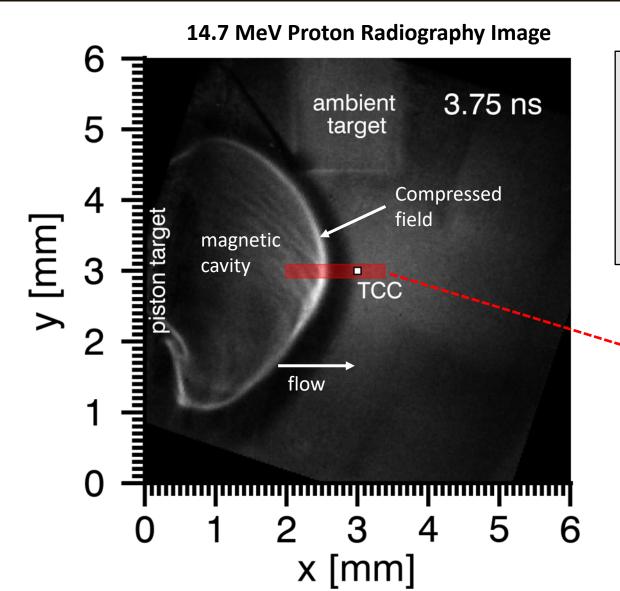
- Interaction probed with Thomson scattering (TS) and proton radiography (PR) diagnostics
- TS probed along expansion direction (x) at TCC
- PR probed the x-y plane by sending protons along z



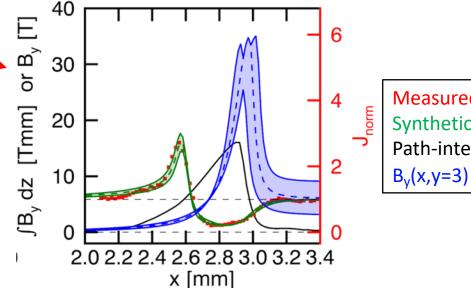
Proton Radiography Indicates Strong Magnetic Field Compression







- Piston plasma sweeps out background field B_y, forming magnetic cavity
- Transition from low fluence (black) to high fluence (white) regions indicate large proton deflections and magnetic compressions
- B_y can be re-constructed by comparing data and synthetic fluence profiles

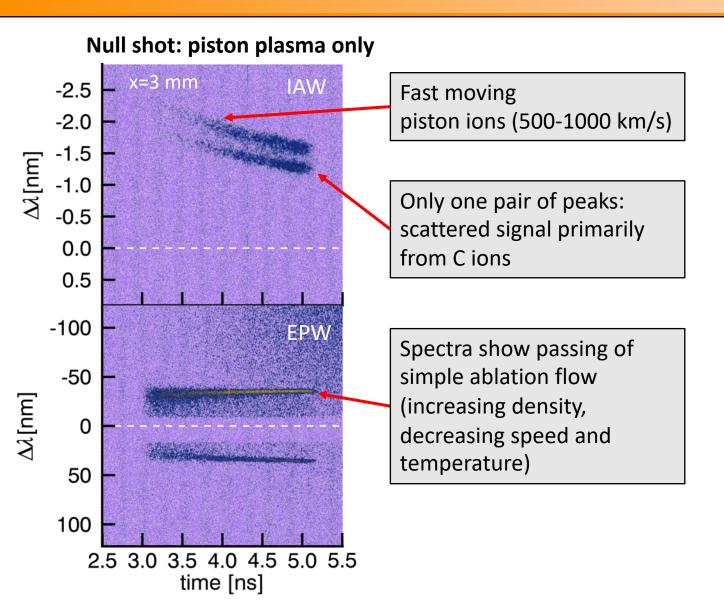


Measured proton fluence Synthetic proton fluence Path-integrated B_y

Spectra Reflect 1D Ion Velocity Distributions



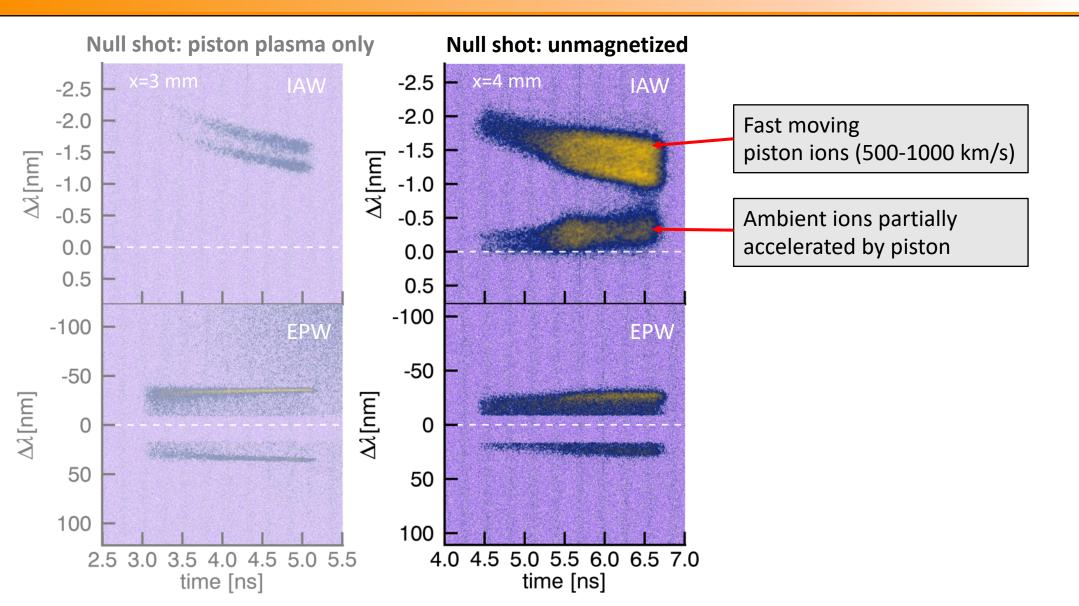




Weak Interaction Observed between Piston and Ambient lons in Unmagnetized System



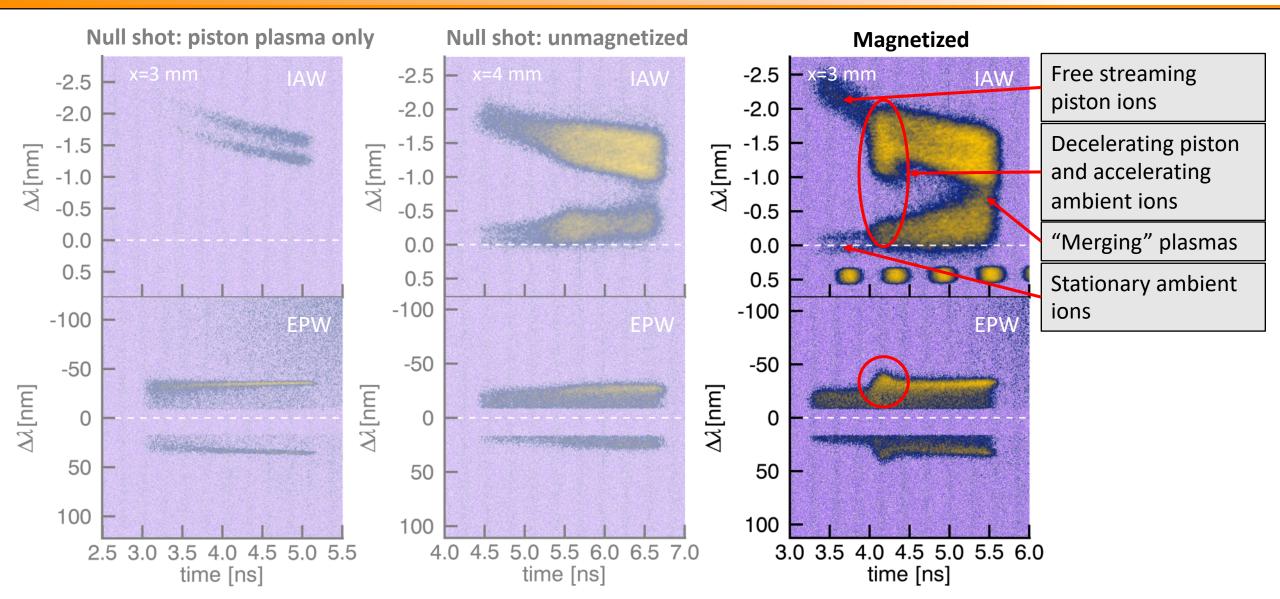




Strong Flow Deformations Observed with Magnetic Field



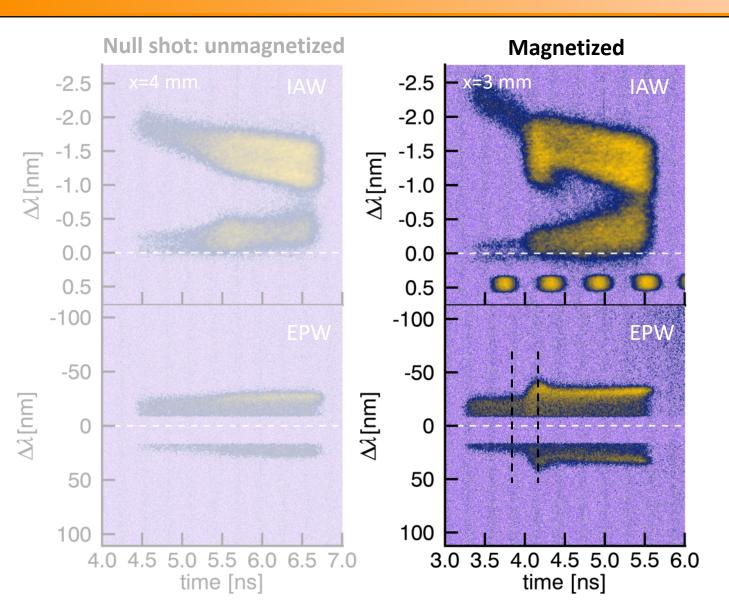


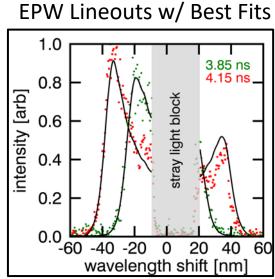


Electron Density, Temperature, and Ion Flow Speed Extracted from Spectra





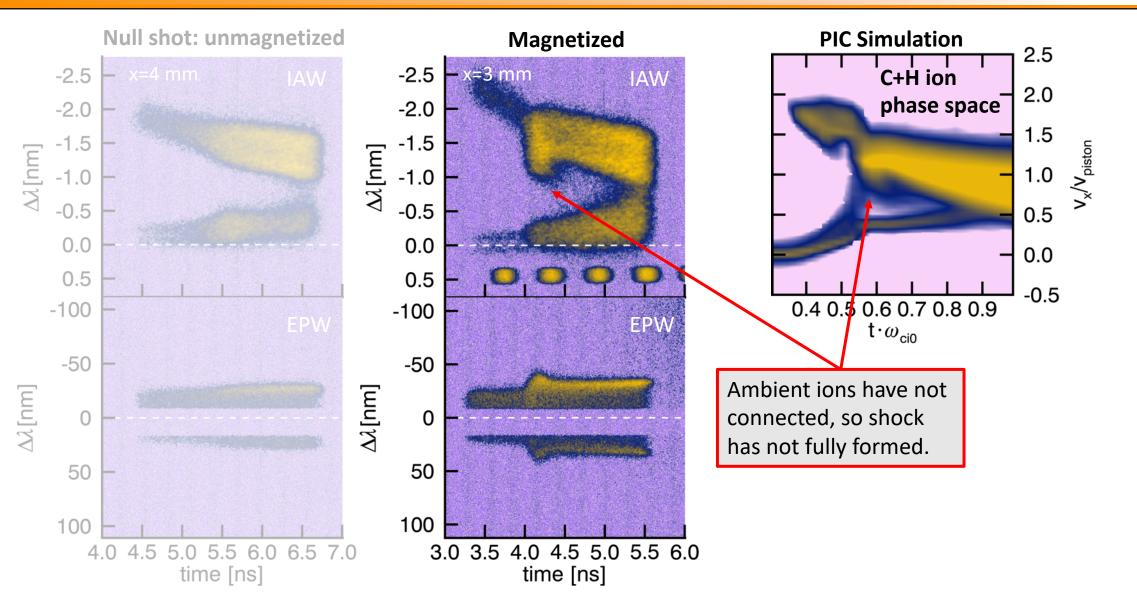




Strong Correspondence between Observations and Simulations



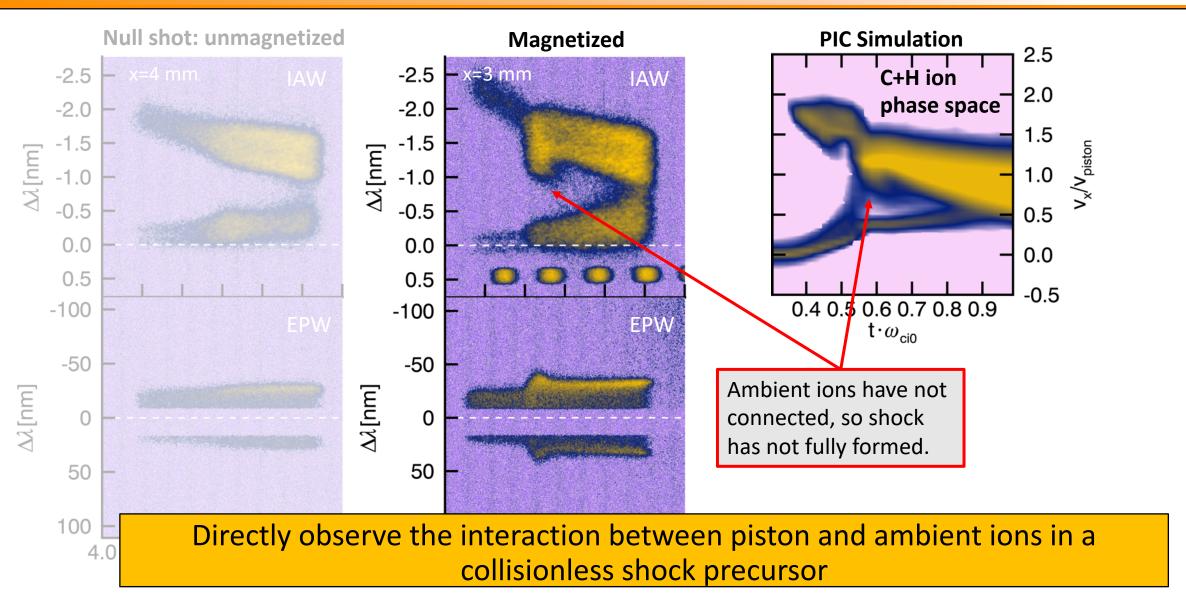




Strong Correspondence between Observations and Simulations



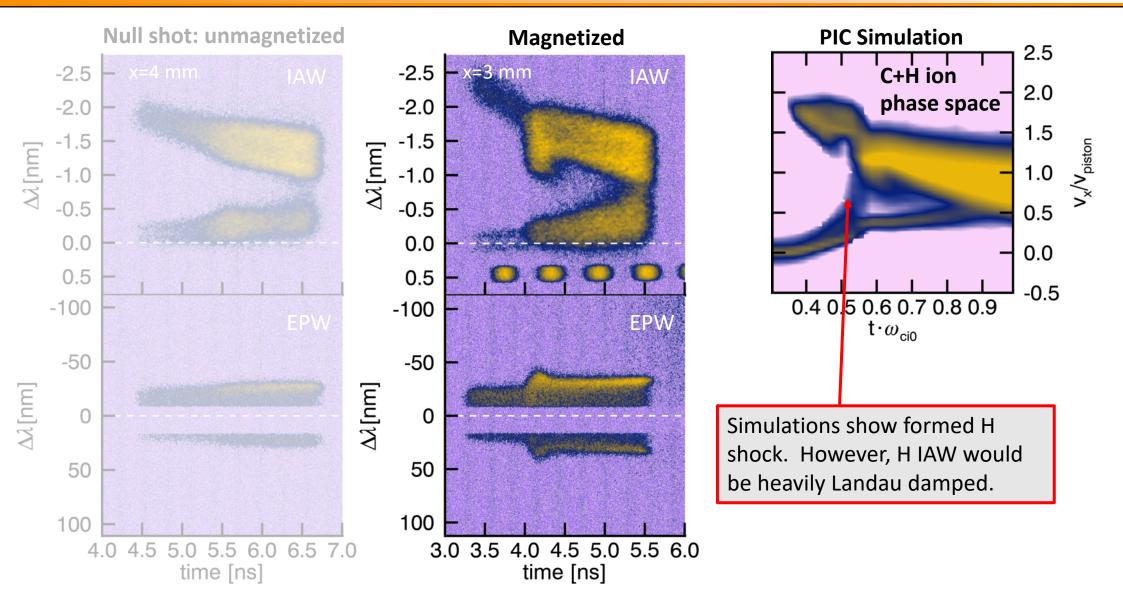




Thomson Spectra May Mask Additional Dynamics in Multi-Species Plasmas

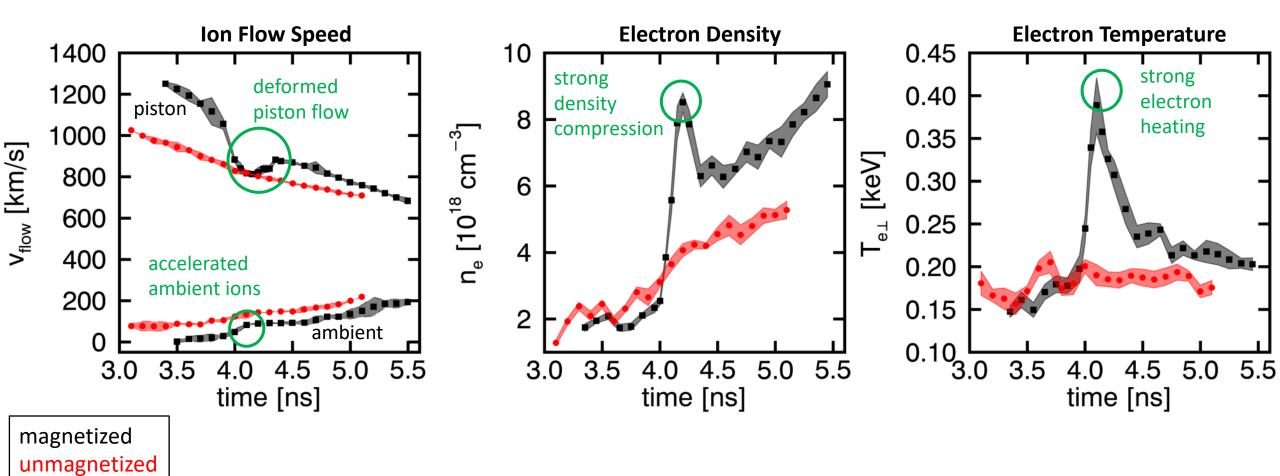






Plasma Parameters Exhibit Strong Dependence on Initial Conditions

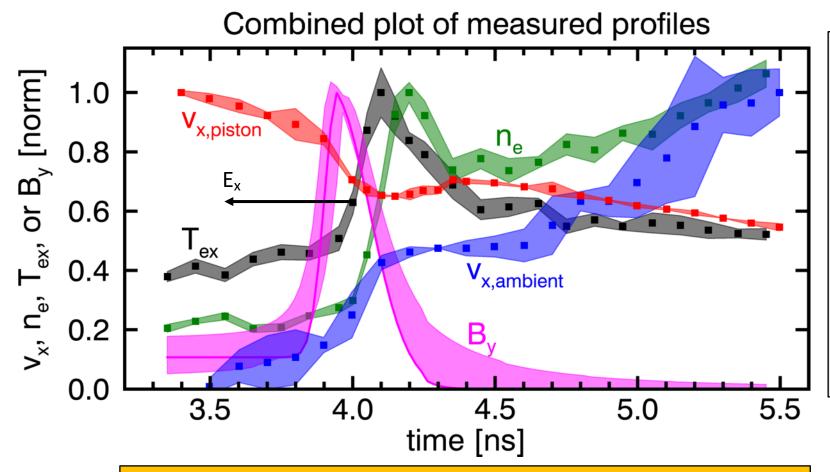




TS and P-RAD Data Show Piston-Ambient Coupling Process







- Magnetic field compressed by ambient plasma
- Piston plasma piles up behind magnetic field
- Electron temperature adiabatically heated by density compression
- Ambient ions accelerated by electron pressure and magnetic field gradients
- Piston flow also modified by these electric fields

Velocity distribution measurements critical to understanding magnetized collisionless shock formation

Outline

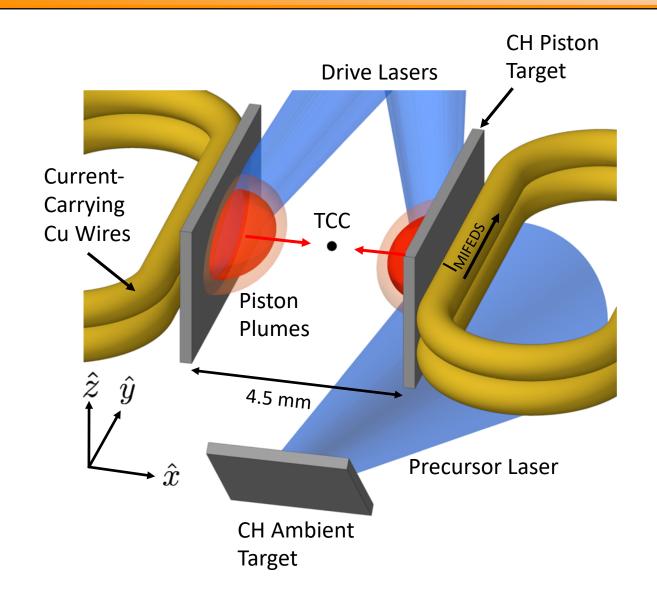


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Experimental Setup for High-M_A Shocks on OMEGA EP





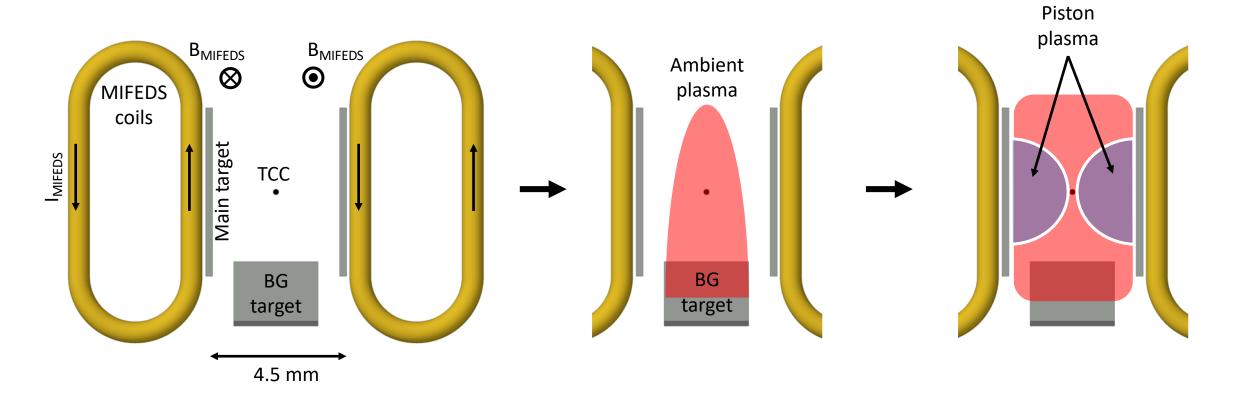


- Interaction probed with angular filter refractometry (AFR), shadowgraphy, and proton radiography (PR) diagnostics
- AFR probed along z
- Shadowgraphy coincident with AFR
- PR probed the x-z plane by sending protons along y

Experimental Setup for High-M_A Shocks on OMEGA EP







MIFEDS coils provide background magnetic field up to 8 T

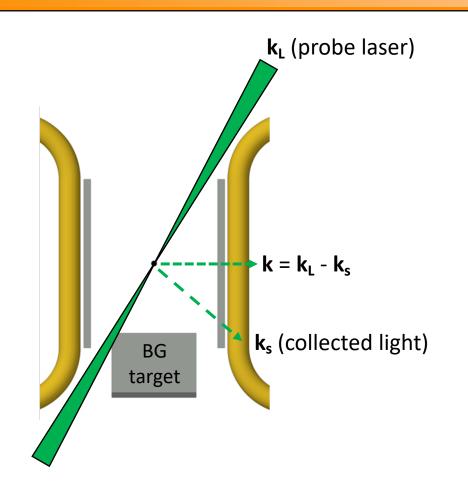
Precursor beam ablates ambient plasma 12 ns before drive beams

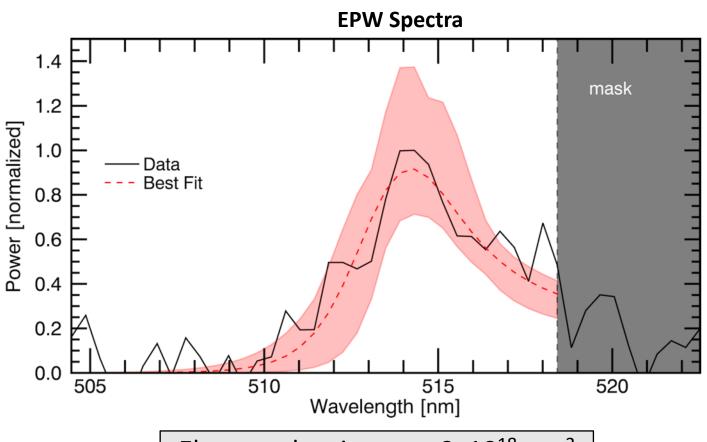
Drive beams create supersonic piston plumes that expand into ambient plasma

Ambient Plasma Characterized with Thomson Scattering









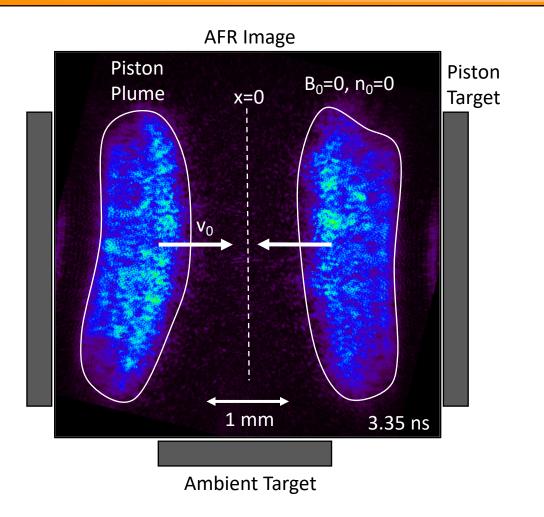
Electron density $n_{e0} = 2x10^{18} \text{ cm}^{-3}$

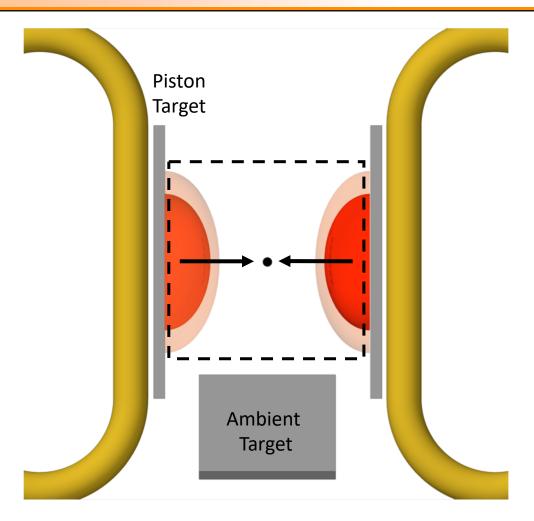
Electron temperature $T_{e0} = 30 \text{ eV}$

Density Evolution Measured with AFR









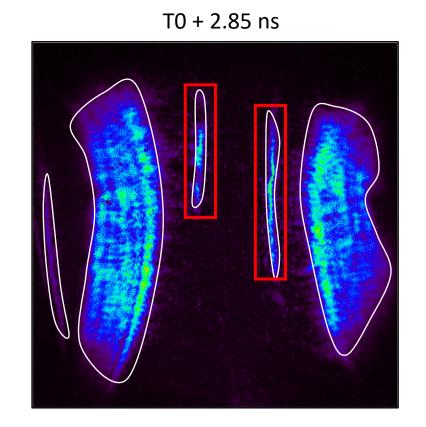
Without background magnetic field or ambient plasma, only piston plumes observed.

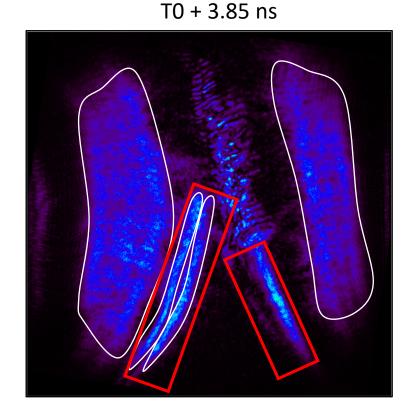
Shock-Like Gradients Observed with B₀>0 and n₀>0





T0 + 2.35 nsPiston Plume Shock-like gradient

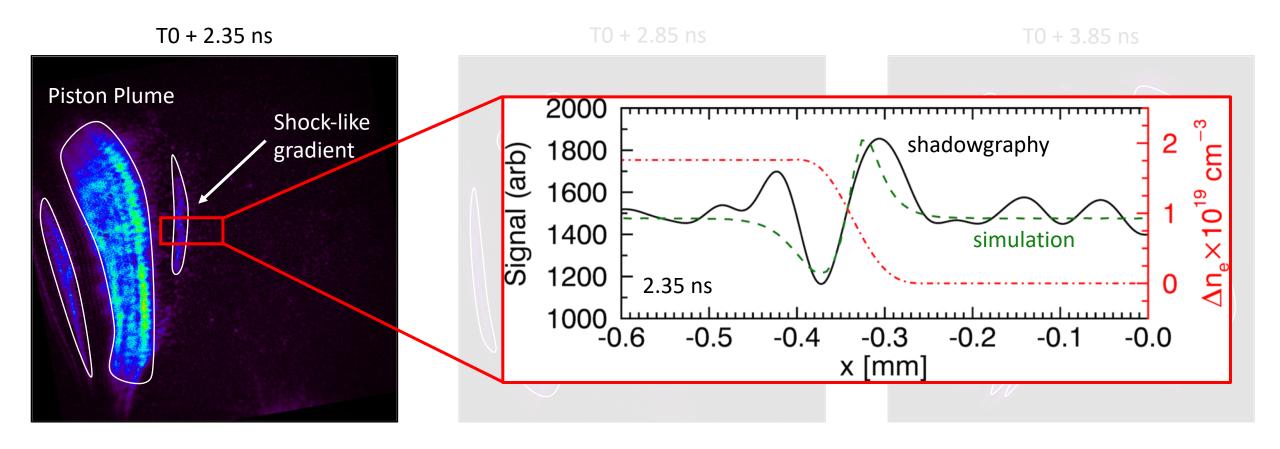




 $\mathbf{v_0} \approx 700 \text{ km/s } (\mathbf{M_A} \approx 15)$

Shock-Like Gradients Observed with B₀>0 and n₀>0



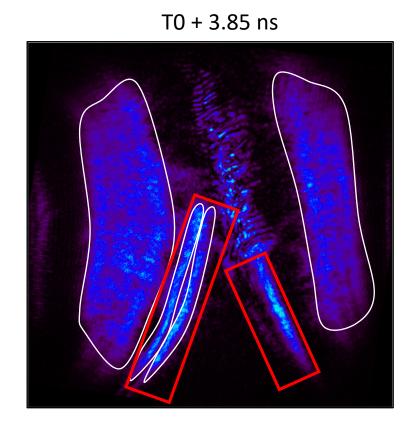


Compression width $\Delta x \sim 0.6 \ c/\omega_{pi}$

Shock-Like Gradients Observed with B₀>0 and n₀>0





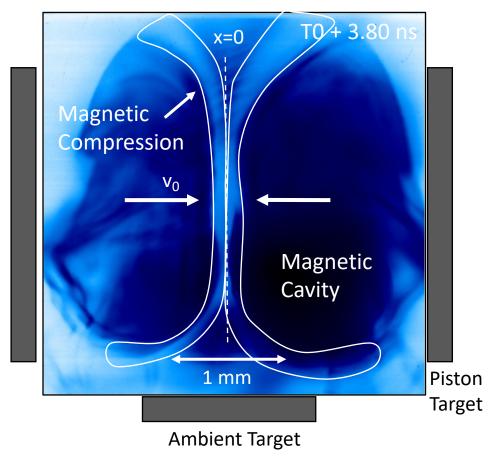


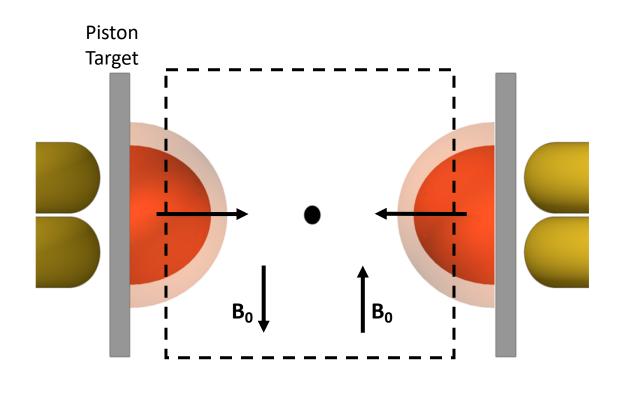
Density compression $n/n_0 > 4$

Magnetic Compressions Observed with Proton Radiography



13 MeV Proton Radiograph





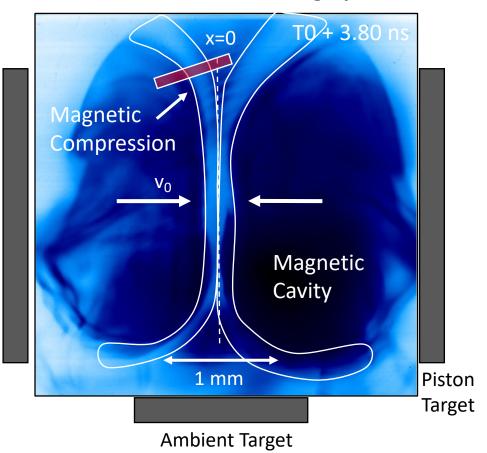
Ambient Target

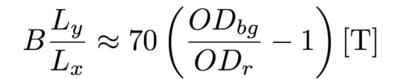
Magnetic Compressions Observed with Proton Radiography

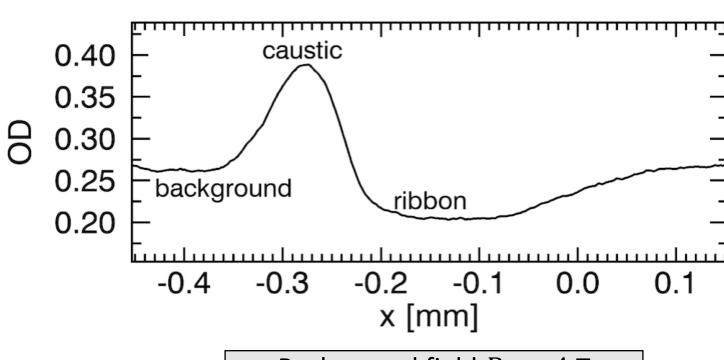










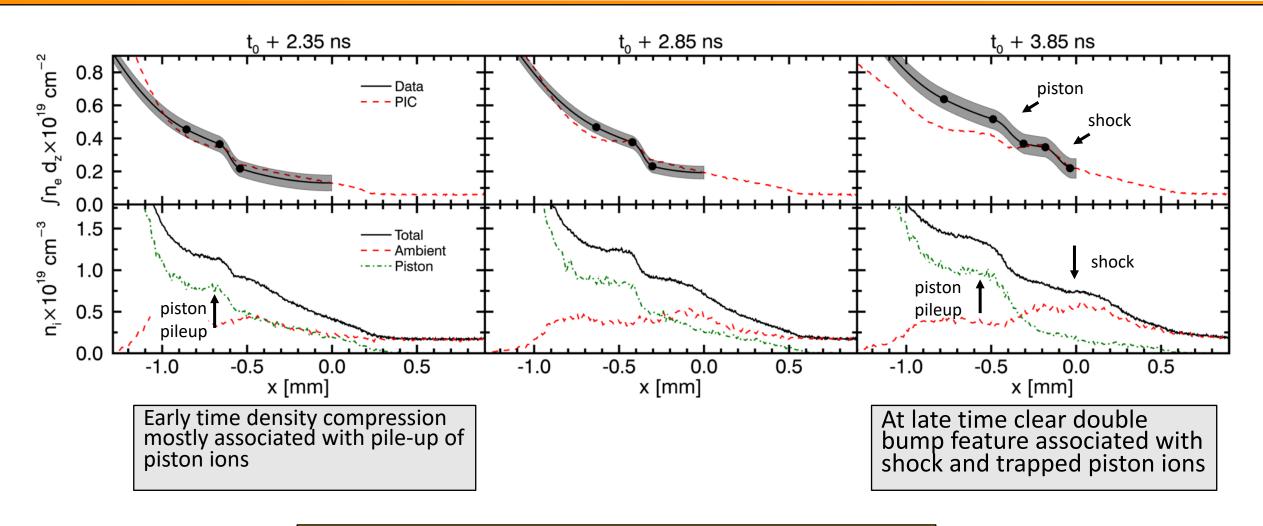


Background field $B_0 \approx 4$ T Magnetic compression $B/B_0 \approx 3$

Density Profiles Show Separation of Shock from Piston





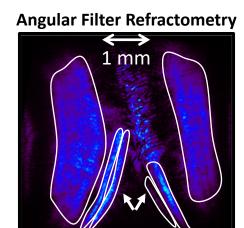


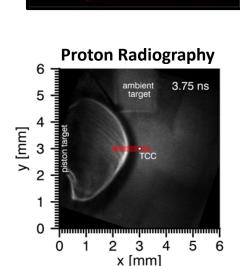
Summary

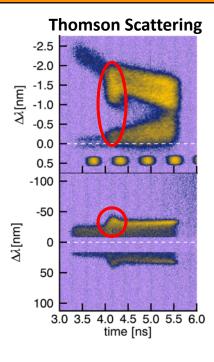


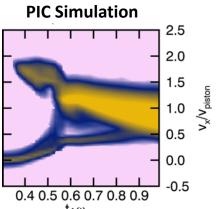


- We have developed a platform for studying laser-driven, high-Mach-number collisionless shocks utilizing advanced diagnostics.
 - First observation of a high-Mach-number magnetized shock in the laboratory
 - First measurements of both ion and electron velocity distributions in a developing magnetized shock
- We have carried out comprehensive particlein-cell simulations that model laser-driven shocks in experimentally-relevant conditions.
 - Observe robust signatures of kinetic-scale shock formation









Future Work



- The development of this platform allows key questions of magnetized shocks to be addressed:
 - Shock heating and energy partitioning
 - Particle injection and acceleration
 - Spatial and temporal scales of shock formation and reformation
 - Interplay between shocks, reconnection, and turbulence
- PIC simulations show that laboratory shocks can lead to non-thermal electron populations
- Experiments underway to explore quasi-parallel collisionless shocks



Thank You!